

Portable turbulence profilers: from MASS to RINGSS

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Outline

- Path to a solid-state upgrade of MASS
- Principle and theory of the RINGSS **method**
- RINGSS **instrument**
- First results at Cerro Tololo and Paranal



New animal in our Zoo!

Turbulence profilers

- Crossed beams: SCIDAR, SLODAR, MCAO, Lunar limb
- Scintillation of Sun/Moon (low layers only)
- Single-star scintillation (MASS, FASS, SHIMM, RINGSS)

Spatial scale: Fresnel radius $\sqrt{\lambda z} = 1.5...10\text{cm}$ (0.5-20km)

Time scale: $20\text{m/s} * 1\text{ms} = 2\text{cm} \Rightarrow$ exp. time 1ms or less

Photon flux: star $\sim 2\text{mag}$, $t_{\text{exp}}=1\text{ms}$, $S=1\text{cm}^2$

Band 100nm, QE 0.5 \Rightarrow 80 electrons

Need fast, low-noise detectors: PMTs, EM CCDs, CMOS

Signal-to-noise vs. aperture in DIMM

- Image motion in DIMM (signal) $\sim D$ (turb. spectrum)
- Centroid photon noise $(\lambda/D) * N_{\text{ph}}^{-0.5} \sim D^{-2}$
- S/N $\sim D^3$!

DIMM uses telescopes of 25-35 cm

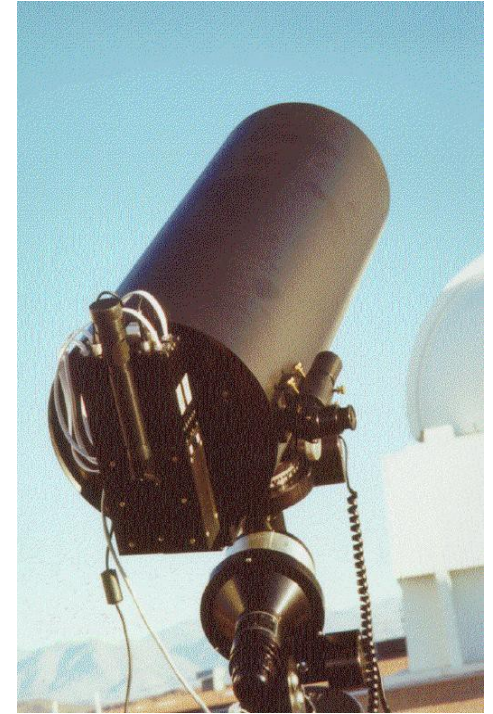
Propagation is important for $d < 10\text{cm}$

Small aperture is a challenge, needs efficient use of photons. Advantage: portability and low cost

The first MASS (1998-2002)



Victor Kornilov
(1953-2021)
Nicolai Shatsky
Olga Vozyakova
Marc Sarazin

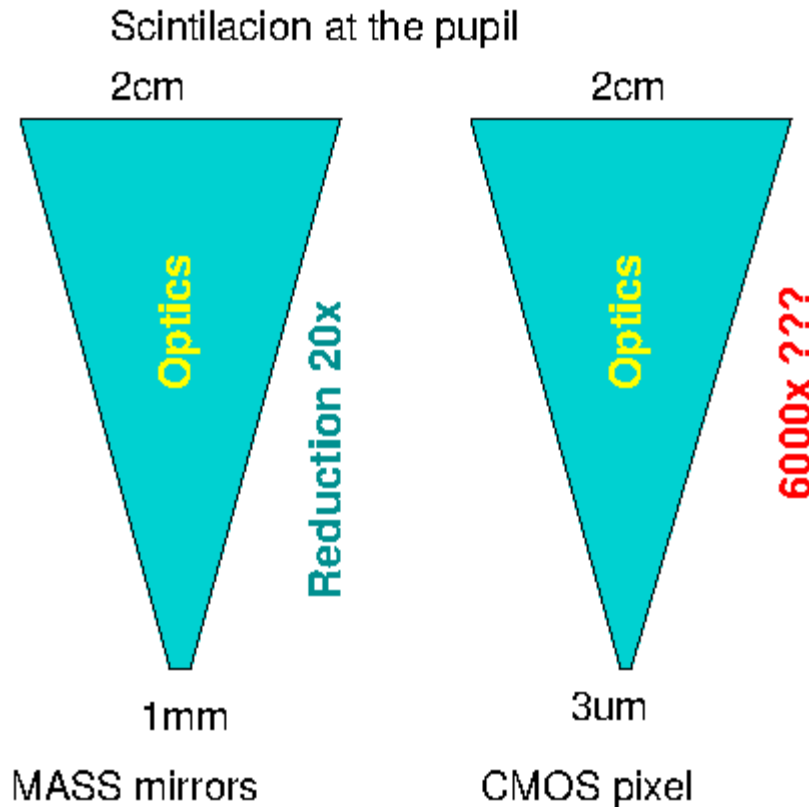


15-cm off-axis reflector
(custom-made)

MASS = theory + instrument + software

How to modernize MASS-DIMM?

- The PMT technology is obsolete → CCD? CMOS?
- Use commercial components (cheaper)

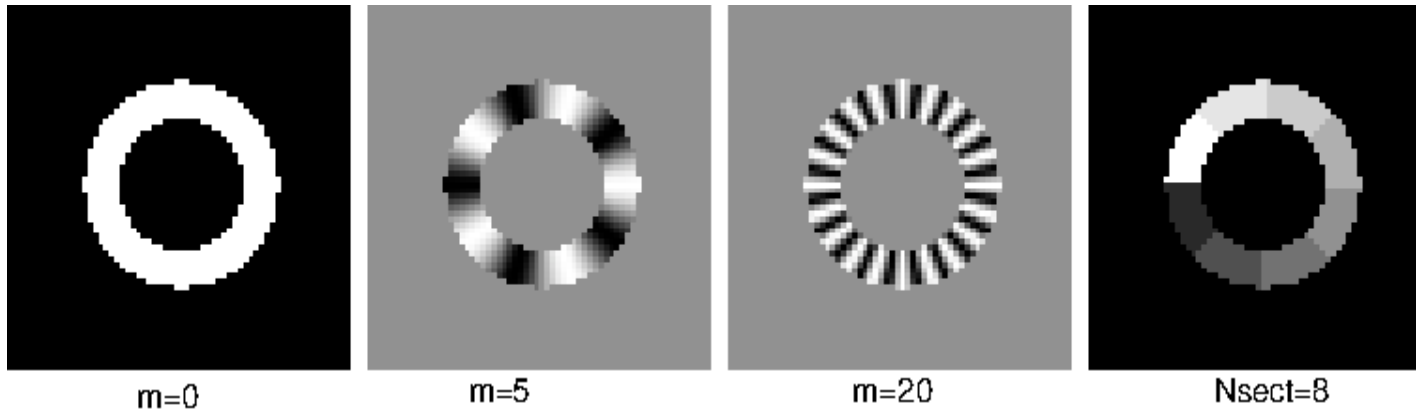


Replacing mirrors
in MASS-DIMM
by a CCD or CMOS
does not work!

It is easier to work in the
image plane than at pupil

Ideas and solutions

- CMOS emulator with binning (LBT prototype, failed)
- EM-CCD with binning (FASS = Full-Aperture Scintillation Sensor), A.Guesalaga (PUC, Santiago)
- Defocused image (ring) → RINGSS
- Fourier transform over angle (FASS), otherwise - aliasing
- SHIMM (Paris, Durham, Moscow, Tokyo): use lenslets

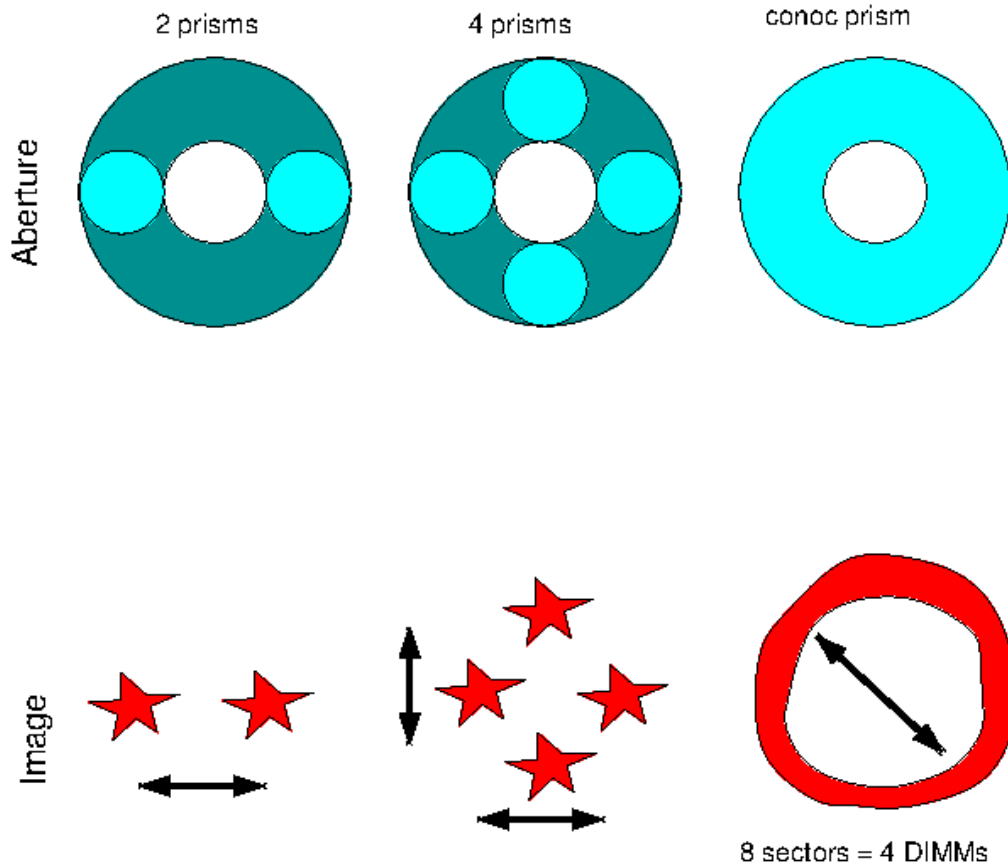


Why ring image?

- Conic wave-fronts are easy to make (defocus+spherical = cone).
- Compared to pure defocus, light is more concentrated (wins over detector noise). But defocus is OK, too.
- Simultaneous seeing measurement by 4 equivalent DIMMs (differential sector motion).
- Rotational symmetry → wind and AO time-constant measurement.
- Use all the light efficiently, can reduce aperture size

Ring seeing monitor has been tested at Tololo in 2007 (A.Kellerer, FADE)

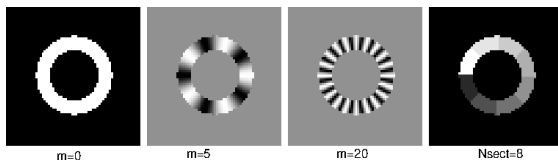
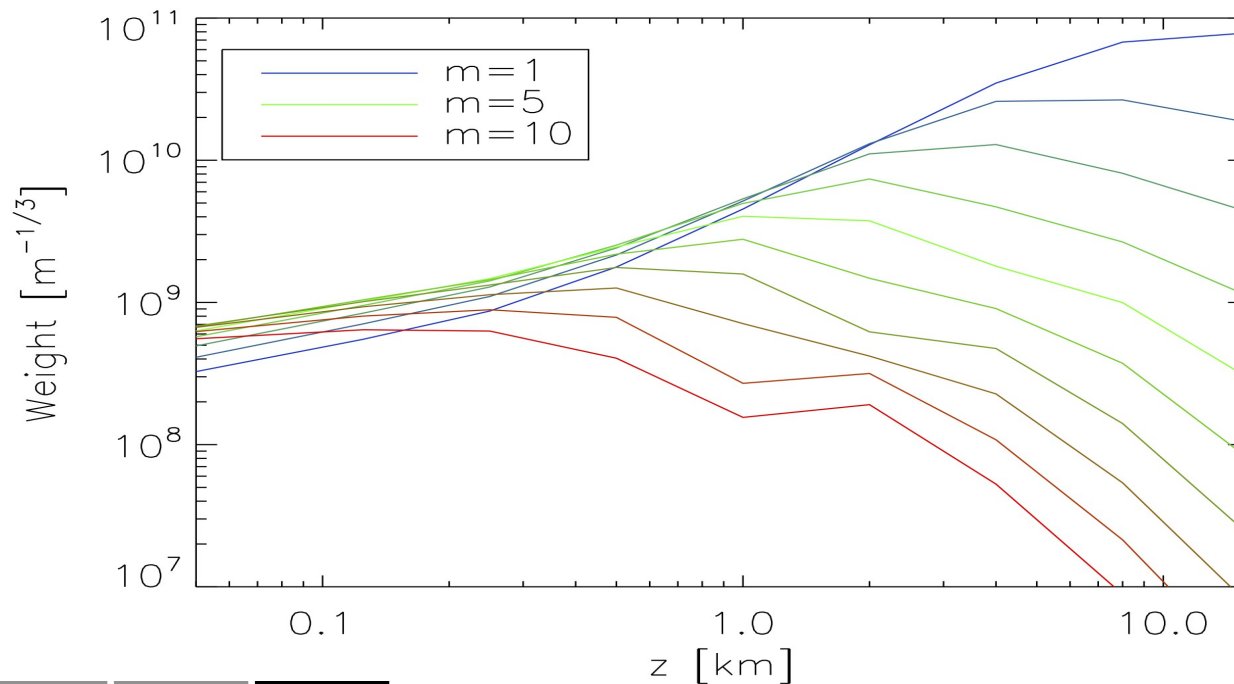
Relation between DIMM and RINGSS



- Conic lens → ring image (replace 2 prisms of DIMM by a conic prism, use full pupil)
- Fluctuations of intensity in the ring serve to measure the scintillation
- Radial distortions of the ring measure the seeing as in DIMM

Theory of RINGSS

- The ring image is not an exact copy of pupil intensity
- Needs modification of the theory to measure turbulence profile.
Sensitive to the ground layer (but also to aberrations)
- Verification of analytic theory and code by simulations



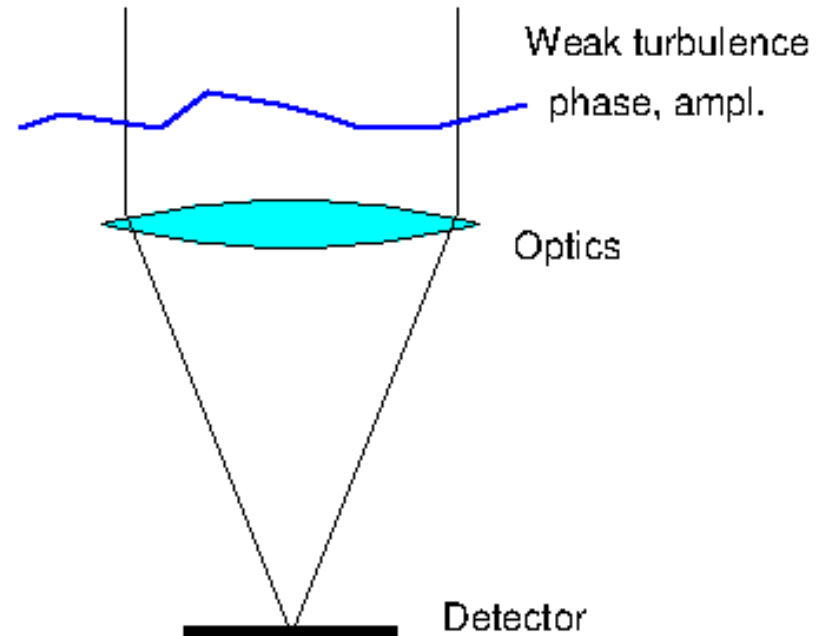
Calculation of weighting functions

2007 MNRAS, 381, 1179

“signal” $S = \text{Image} * \text{Mask}$

$$\Delta S = \int F_{\varphi} \varphi + \int F_{\chi} \chi$$

$$\langle \Delta S^2 \rangle \rightarrow |F_{\varphi} + F_{\chi}|^2 \rightarrow \text{WF}$$



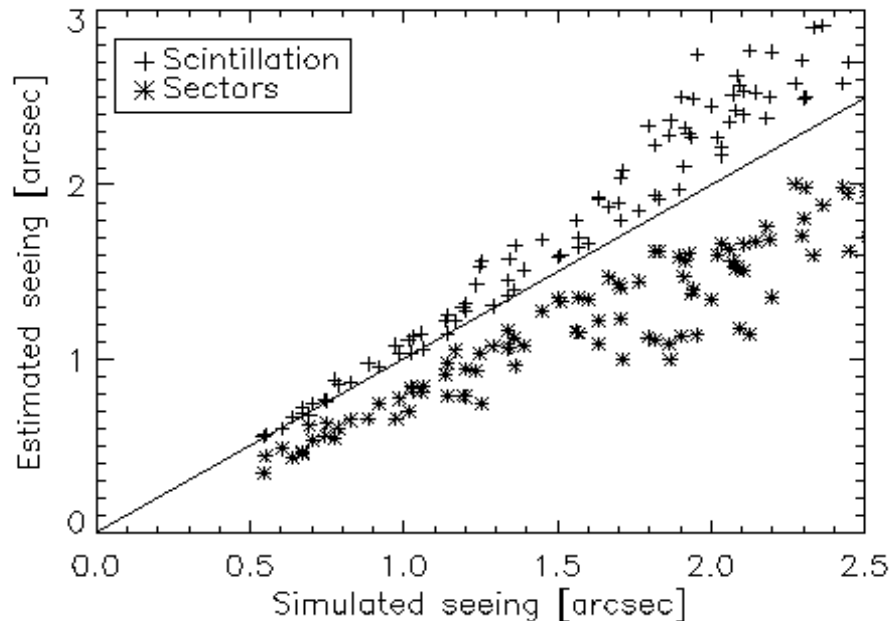
DIMM: centroid mask \rightarrow reaction to defocus etc.

RINGSS: $\sin/\cos(m\theta)$ mask + conic \rightarrow WF(m)

FASS: $\sin/\cos(m\theta)$ mask + defocus \rightarrow WF(m)

Biases

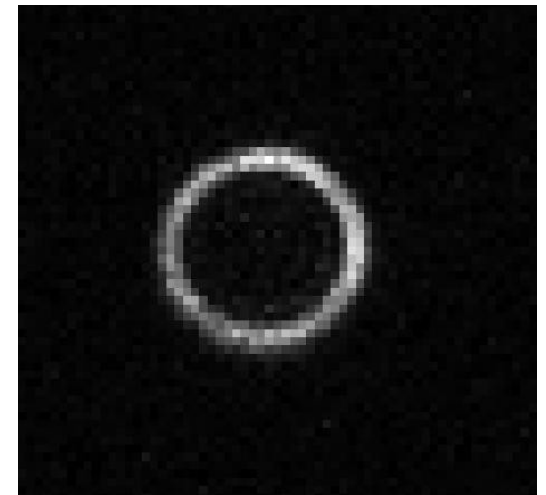
- Noise: calibration (e/ADU) and analytic calculation
- Exposure time: extrapolate to $t_{\text{exp}}=0$ using covariance
- Saturation: simulation-based correction of power spectrum (use machine-learning instead?)
- Aberrations: measure from ring and modify the WFs.



These biases must be accounted for in all single-star profilers

The advantages of RINGSS

- Commercial components: easy to replicate (except enclosure), modern technology (CMOS)
- Telescope smaller than in MASS-DIMM
- Wide field, easier to point the star
- Better data (QE, more pixels, 4 DIMMs)



RINGSS is being replicated in Australia and at LBT.
The signal-processing software is publicly available.

The instrument

Telescope Celestron 5SE

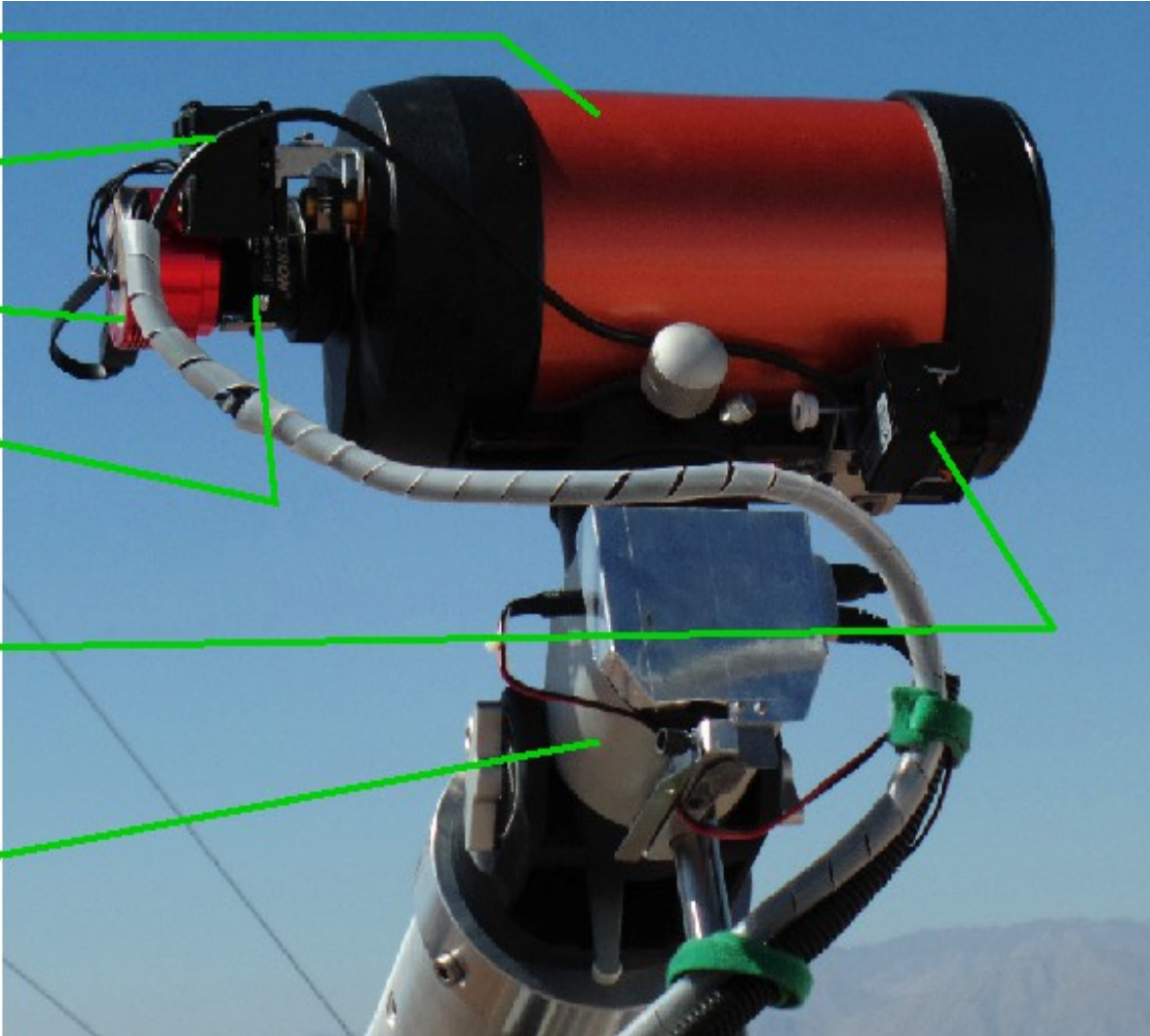
Focuser

Camera ZWO ASI290MM

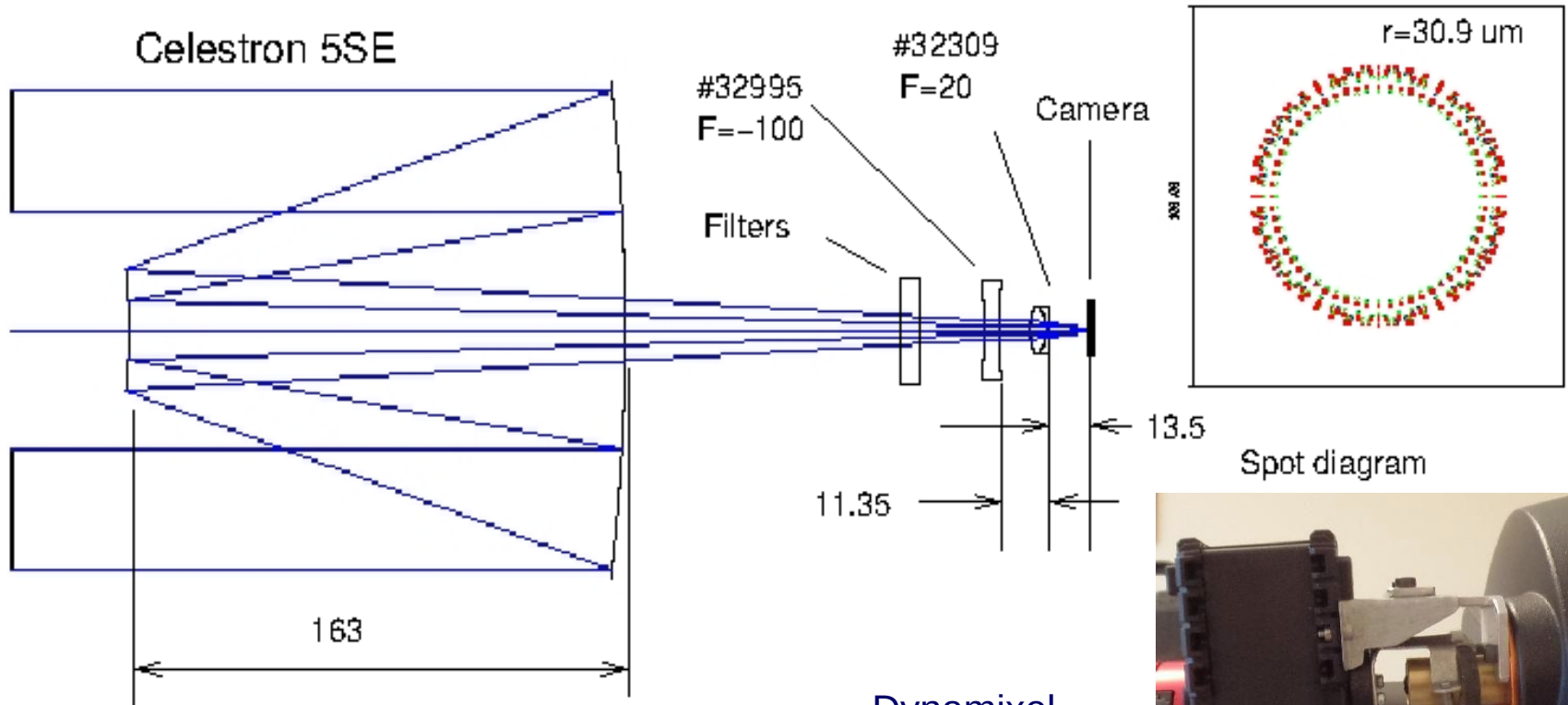
Lenses and filters

Pointing camera

Mount RST-135

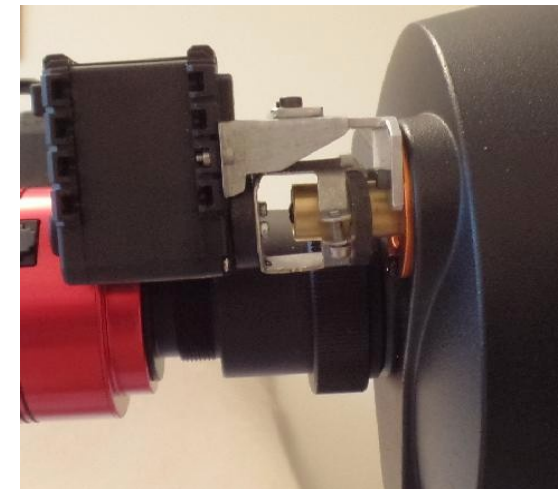


Optics and focuser



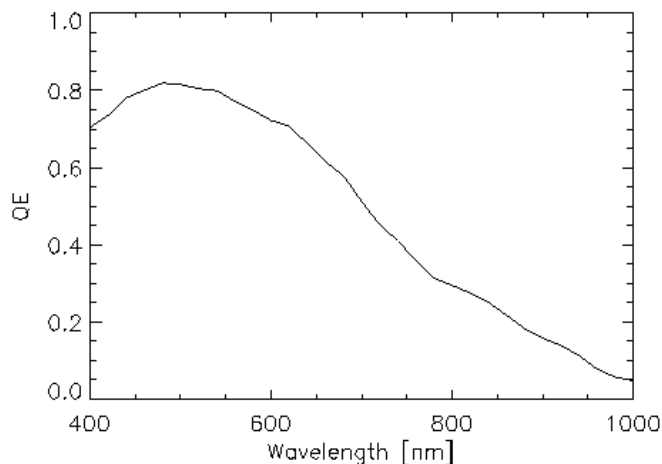
Pixel scale 1.7"
Ring radius 10 pixels
Full field ~ 0.5 deg

Dynamixel
AX-12A
attached to
the focus
knob



The CMOS detector

- ZWO ASI 290MM (China) used in astro-photography
- Format 2Kx1K pixels, 2.9 $\mu\text{m}/\text{pixel}$
- Quantum efficiency $\sim 80\%$, noise 1 electron
- Interface USB, rapid enough
- Cost $< \$400$ USD!
- Studied in Feb. 2020 (L.Peige, M.Bonati, B.Cancino)



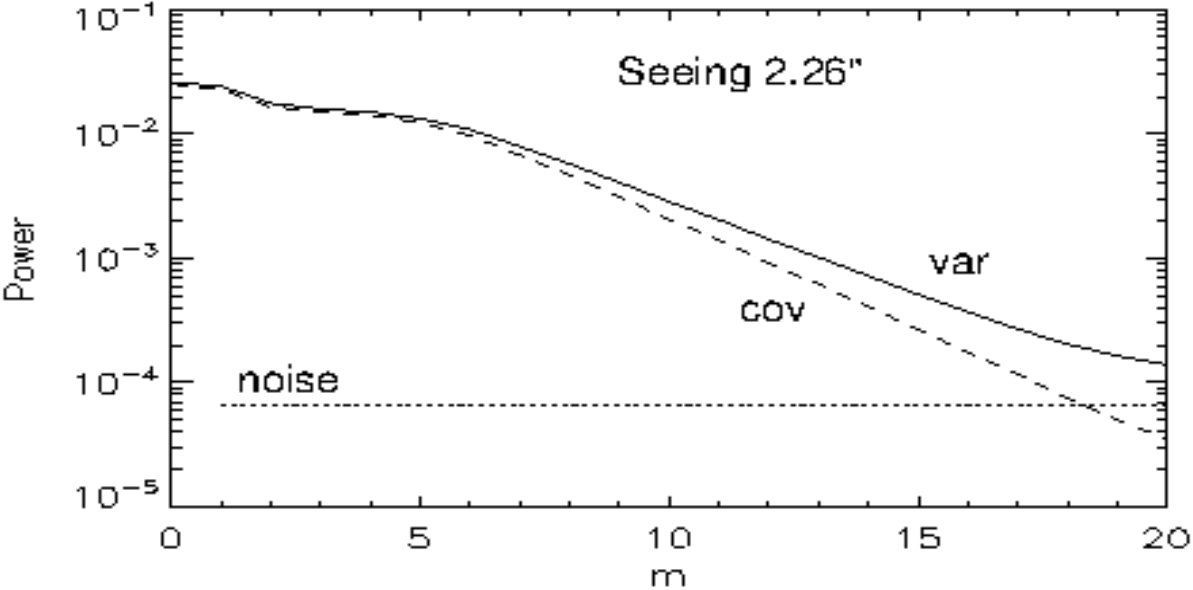
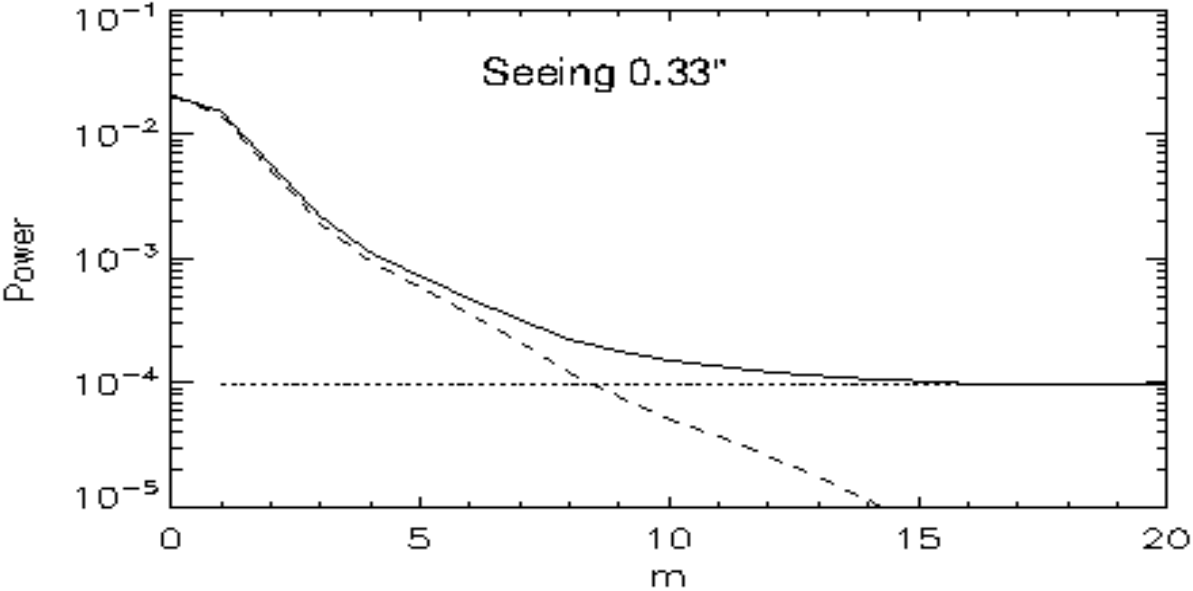
Classic CCD is too slow
and noisy, EM CCD
is expensive ($>25\text{K}$ USD)

Signal processing

- Determine ring radius and width
- Prepare the sine/cosine/sector masks up to $m=20$
- Multiply each frame of the cube by masks → “signals”
- Signal variance → angular spectrum and sector motion
- Average spectrum over ten 2-s data cubes
- Restore profile using pre-computed polychromatic weighting functions (star color matters)
- Measure wind speed and AO time constant (Kornilov 2011)

Signal processing is coded in IDL and python, extensively tested.
Code posted at <http://www.ctio.noirlab.edu/~atokovin/ringss/>

Angular power spectrum



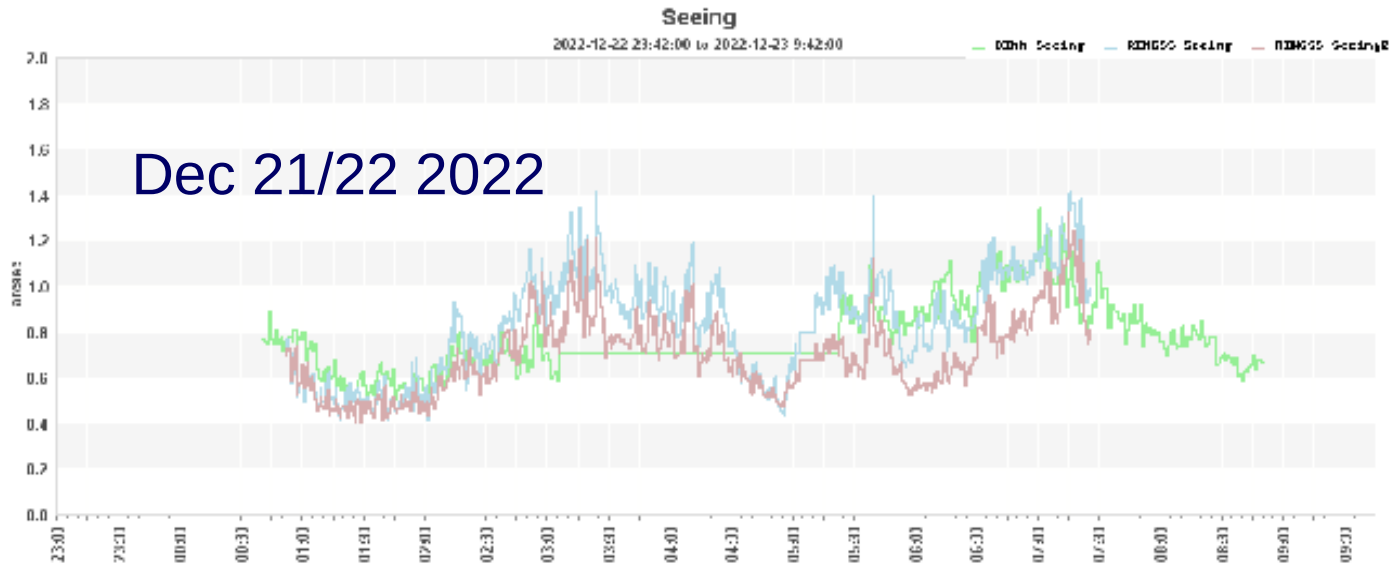
Paranal,
March 2023

RINGSS-1 on the Halfmann tower

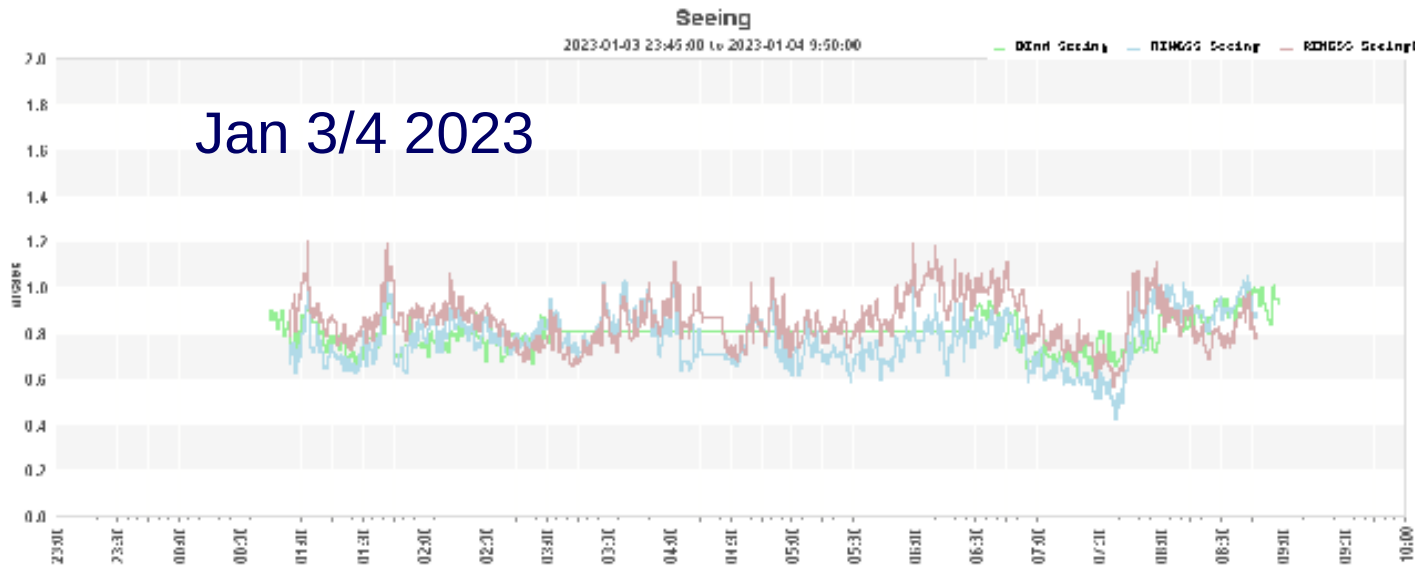


RINGSS on the Halfmann tower at Cerro Tololo
October 2022

RINGSS-1 on duty (robotic)



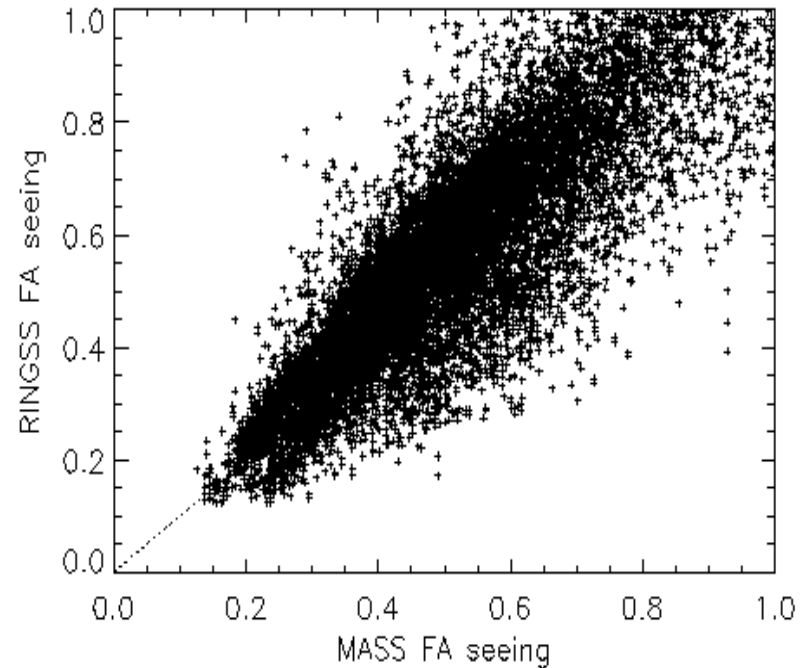
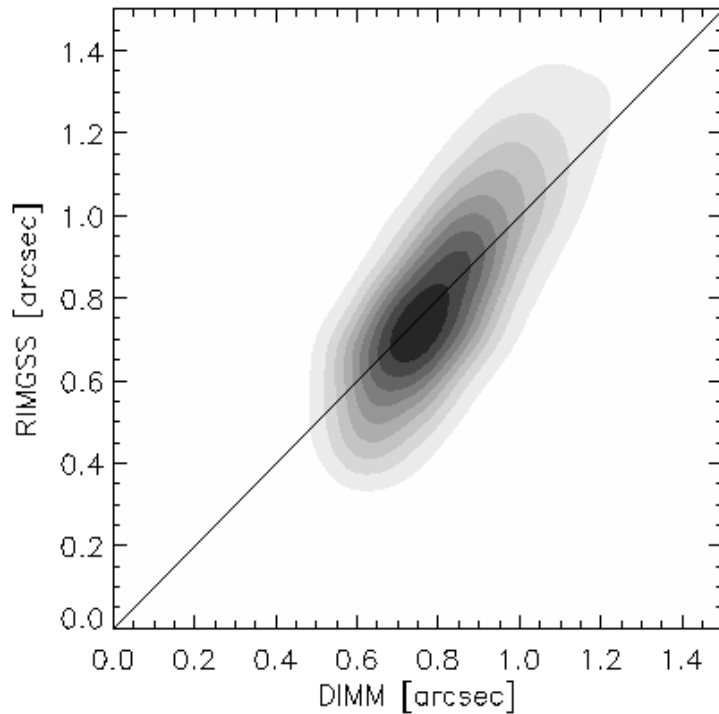
green: DIMM
blue: scintillation
red: sectors



COAT March 30, 2023

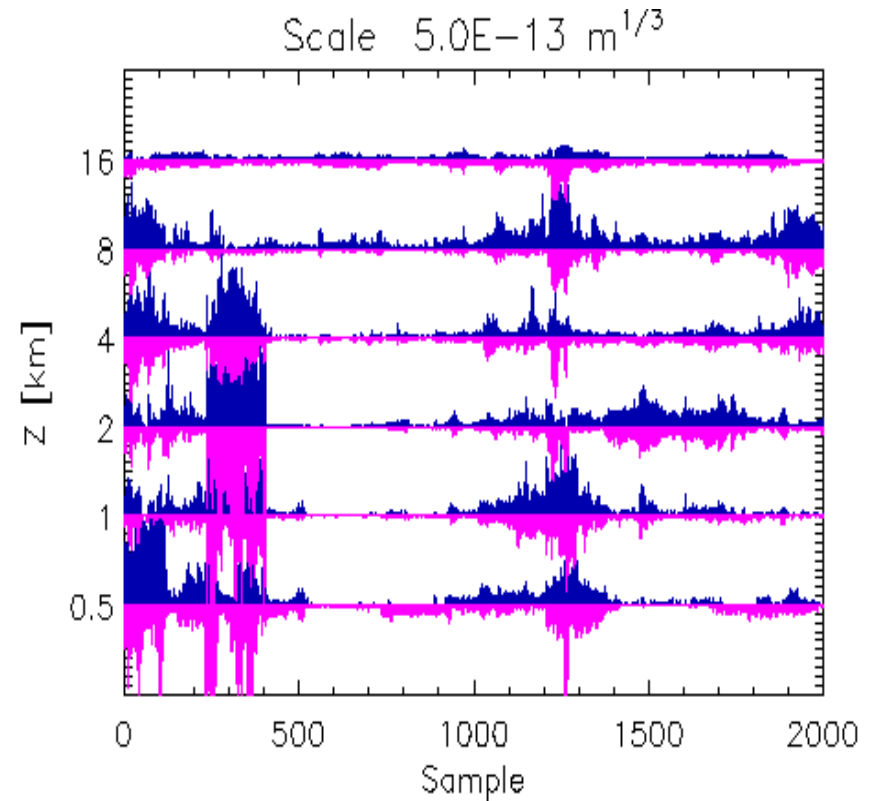
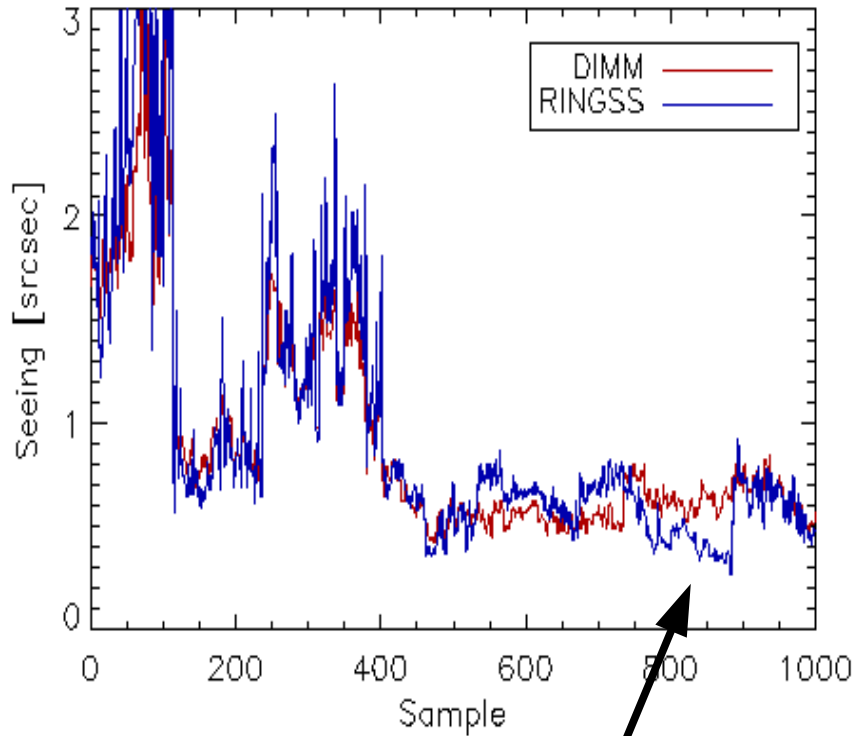
RINGSS vs. MASS-DIMM at Tololo (1)

- Nov. 2022-Feb. 2023: 18,900 matched measurements
- Full report: on the RINGSS web page



$$\text{RINGSS} = -0.27 + 1.36 * \text{DIMM}$$

RINGSS vs. MASS-DIMM (2)



“Good seeing” problem:
DIMM over-shoots

RINGSS: blue-up bars
MASS: pink-down bars

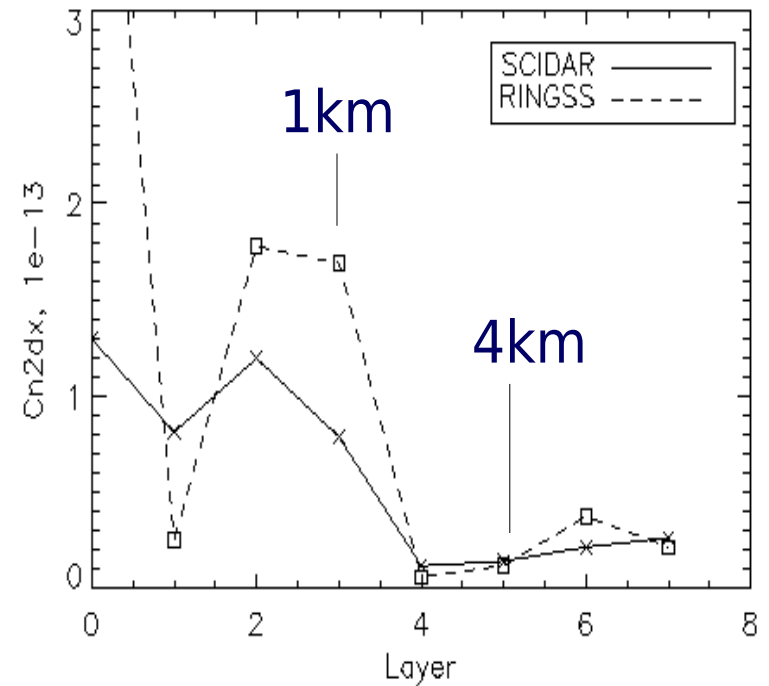
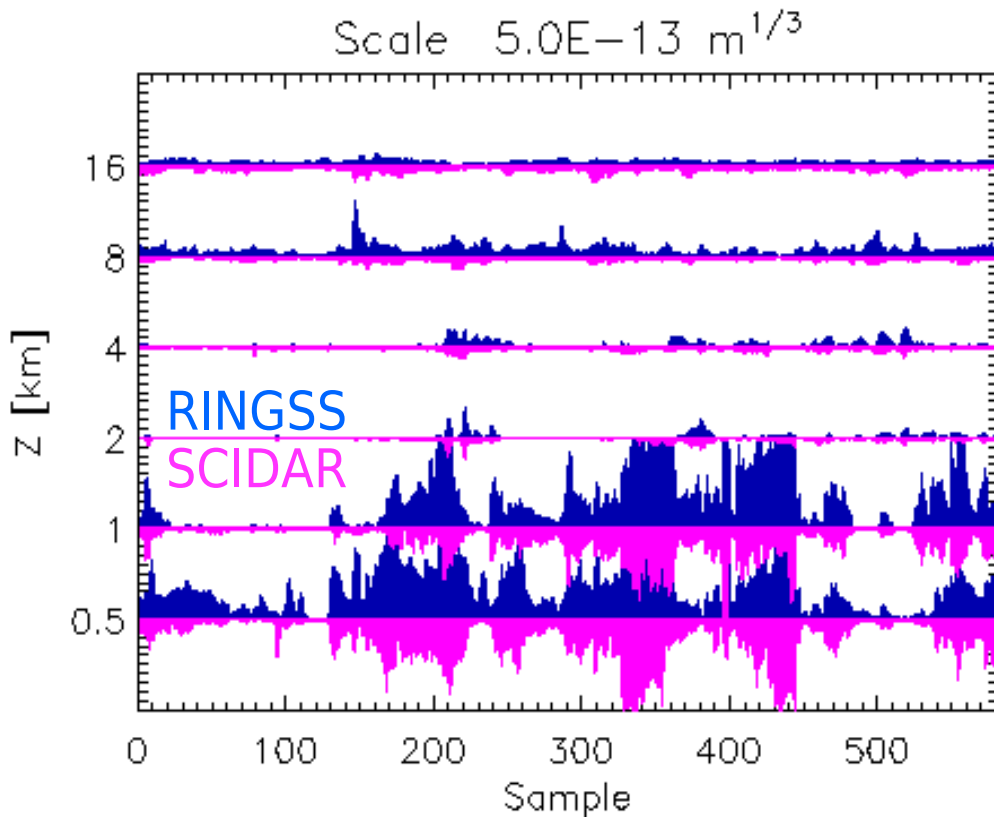
RINGSS-2 at Paranal

- Campaign Feb 28 – March 4 organized by ESO
- SHIMM, RINGSS, FASS plus MASS-DIMM and SCIDAR



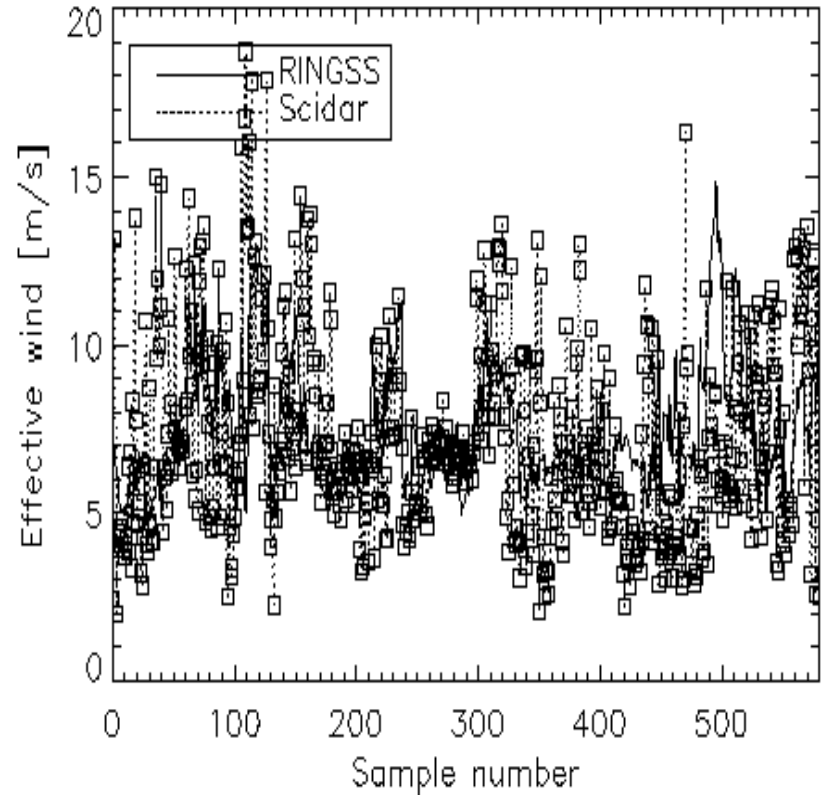
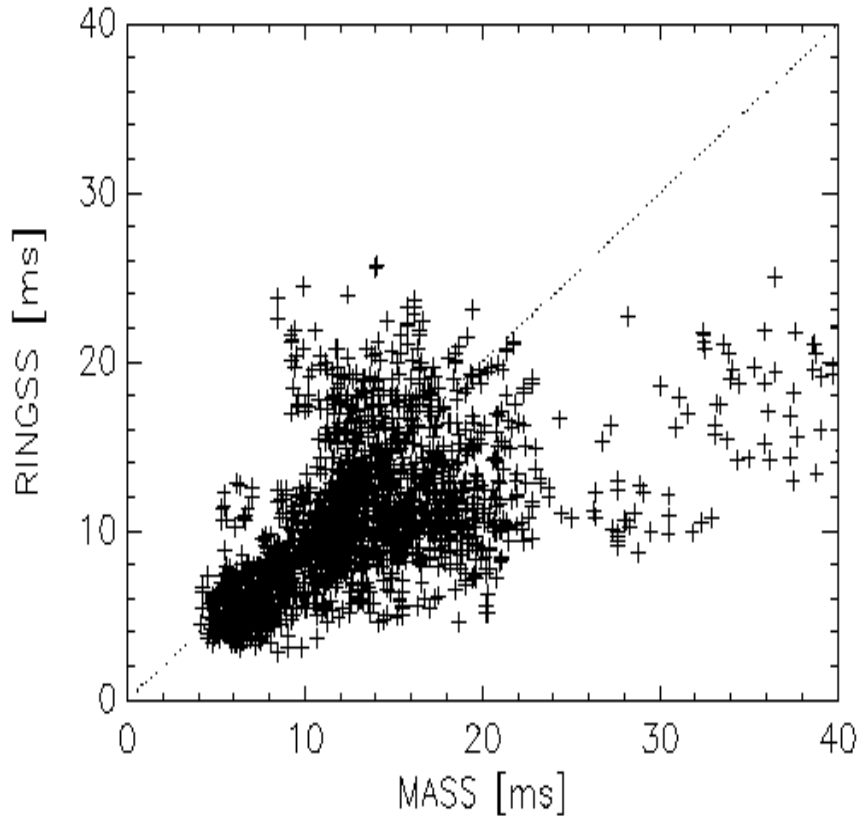
RINGSS vs. SCIDAR

● Credit: T. Butterley (data proc.), J. Velasquez (obs.): 611 profiles with 2-min. resolution during 5 nights.



Mean prof.

Wind and time constant



SCIDAR: two methods to compute V_0 .

Compact enclosures for RINGSS

- Cylindrical equatorial shell (Tololo)
- Compact box for Alt-AZ (simpler, home-made)
- No enclosure: protect only optics and mount?



RINGSS fits in
2 suitcases
(without enclosure)

Summary

- The RINGSS concept is new but mature, anchored in the MASS experience and theory
- The instrument works in robotic mode and gives correct results
- Ready for replication or modification.

RINGSS **method** and data processing: public
RINGSS **instrument** operation: proprietary