

Near Infrared Photometric Parameters of Bulge Globular Clusters from the VVV Survey

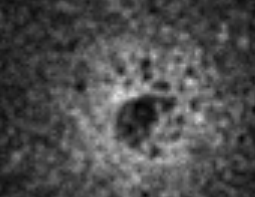
Roger Cohen (U. de Concepción)

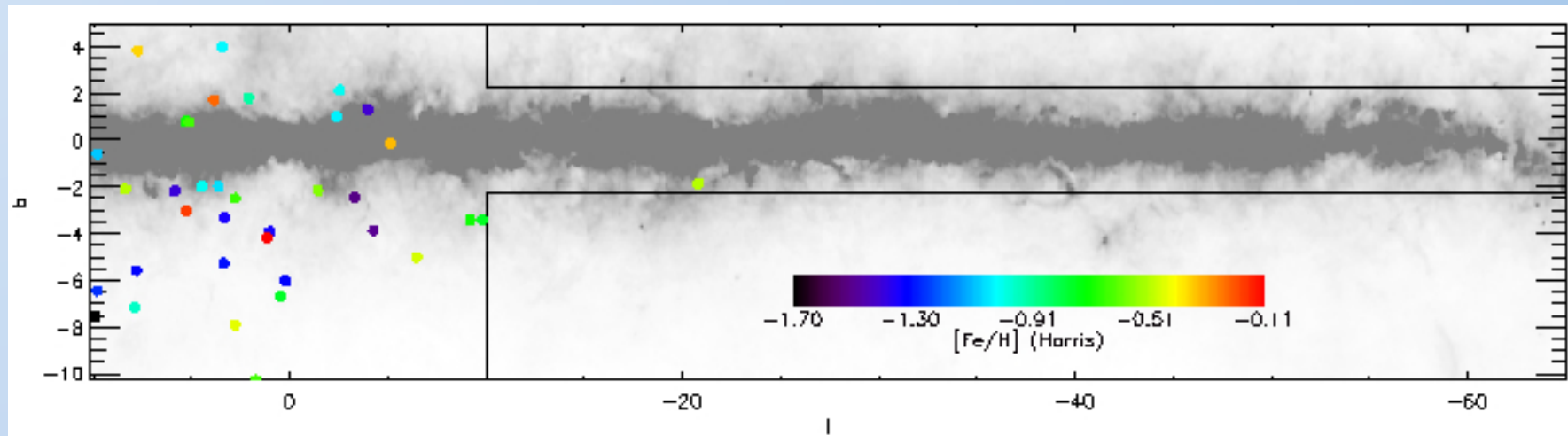
CTIO 50 Years: La Serena, Chile, May, 2013

Co-Workers Include:

C. Moni Bidin
F. Mauro
A.-N. Chene
J. Cummings
S. Villanova
D. Geisler

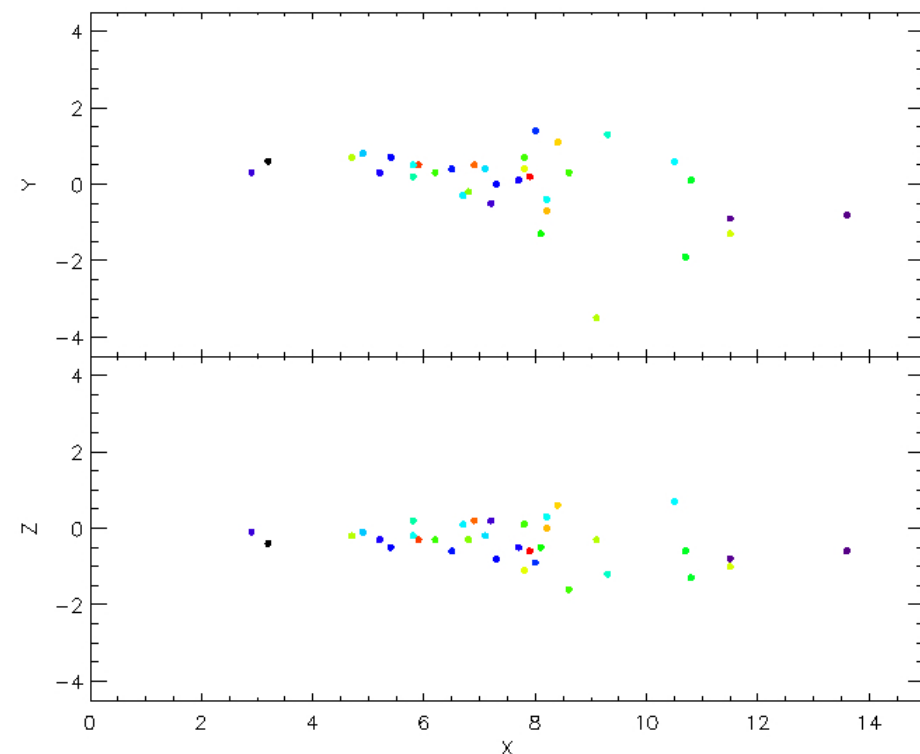
J. Alonso
M. Catelan
M. Hempel
C. Bonatto
J. Vanderbeke
and more!

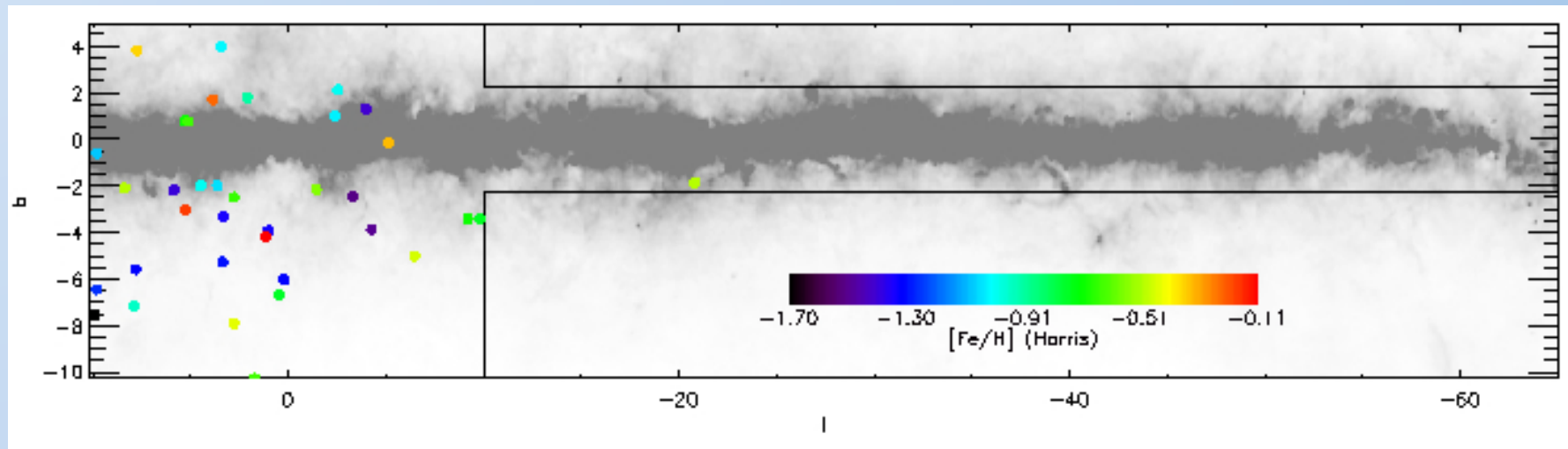




Bulge GGC formation/evolution is an open issue! (AMR)

- Only 17/65 inside 4 kpc have *any* systematic age measured
De Angeli+05, Marin-Franch+09, Dotter+10
- Hi-res abundances $[Fe/H]$, Call
Carretta+09, Saviane+12, more coming!





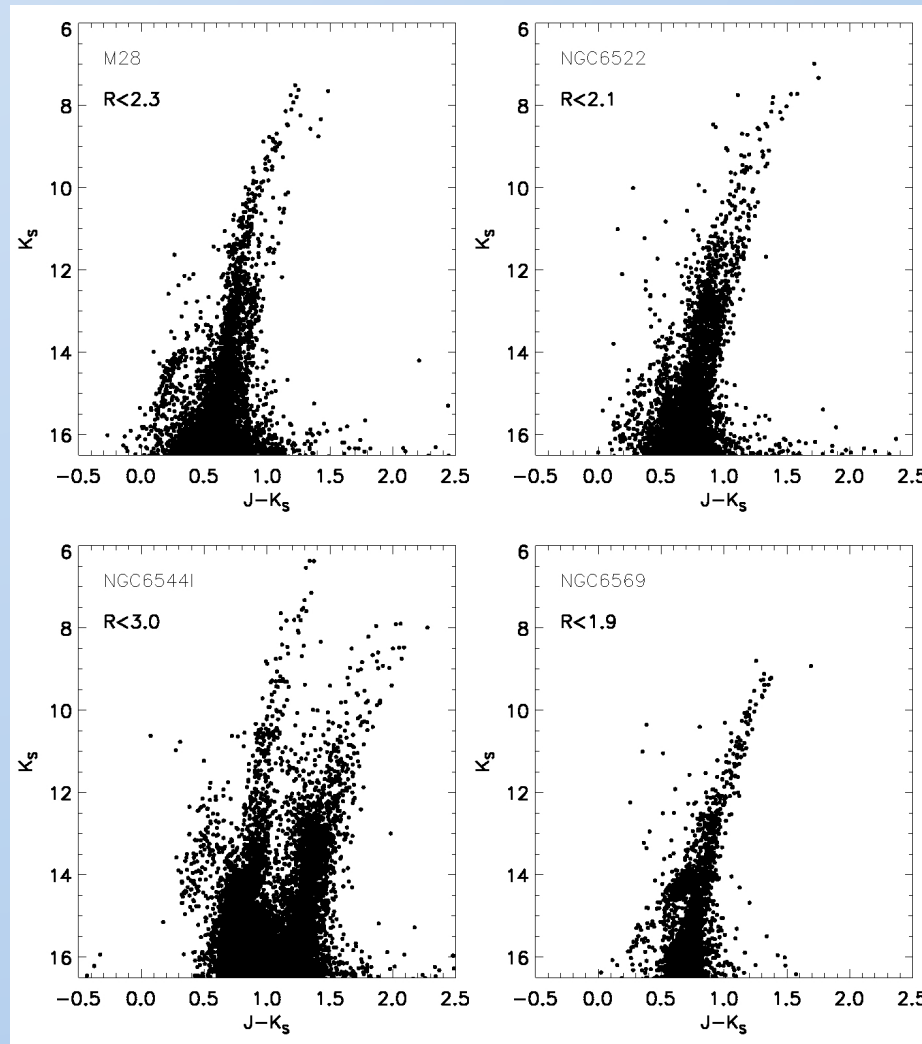
Bulge GGC formation/evolution
is an open issue! (AMR)

- Only 17/65 inside 4 kpc have
any systematic age measured
De Angeli+05, Marin-Franch+09, Dotter+10
- Hi-res abundances $[\text{Fe}/\text{H}]$, Call
Carretta+09, Saviane+12, more coming!

VVV!

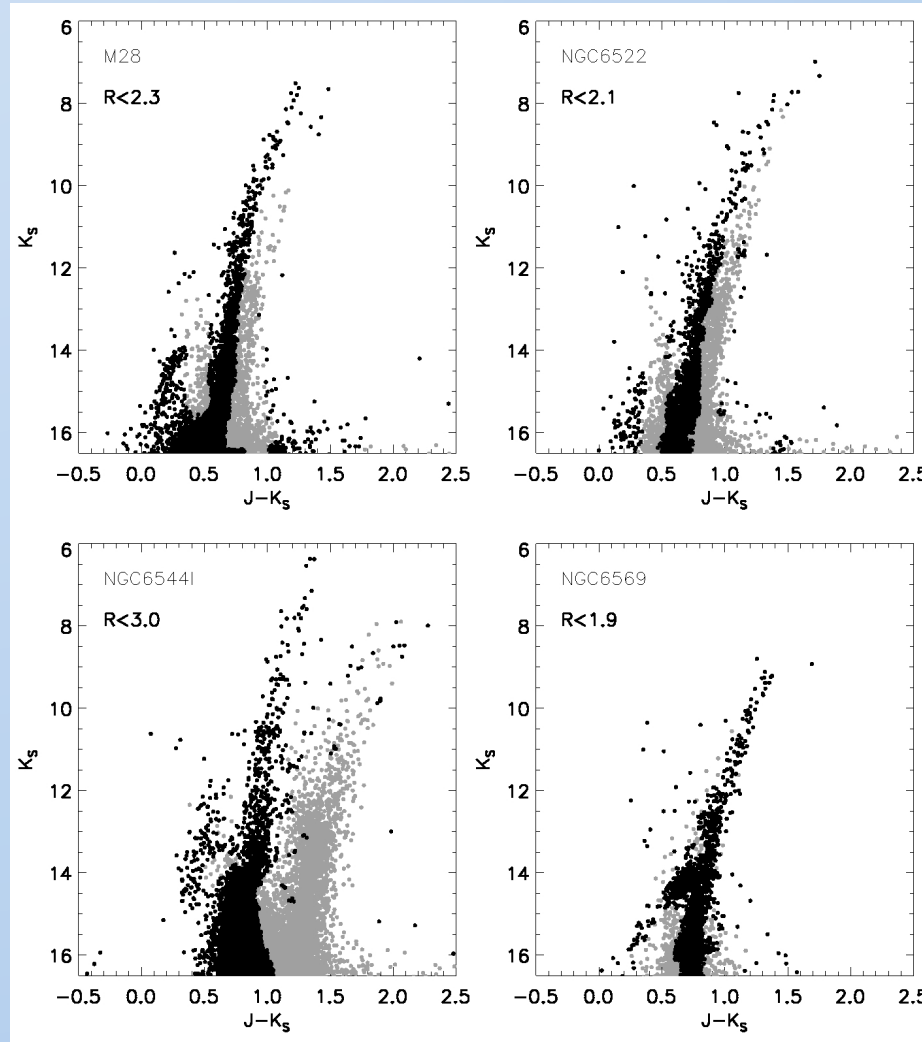
- ~36 GGCs in VVV survey area
- ~16 have spectroscopic abundances
- PSF photometry, calibrated to 2MASS
(VVV-SkZ pipeline, *Mauro+13*)
photometric zpt uncertainty < 0.02 mag,
astrometric RMS ~0.2"

Red Giant Branch Analysis



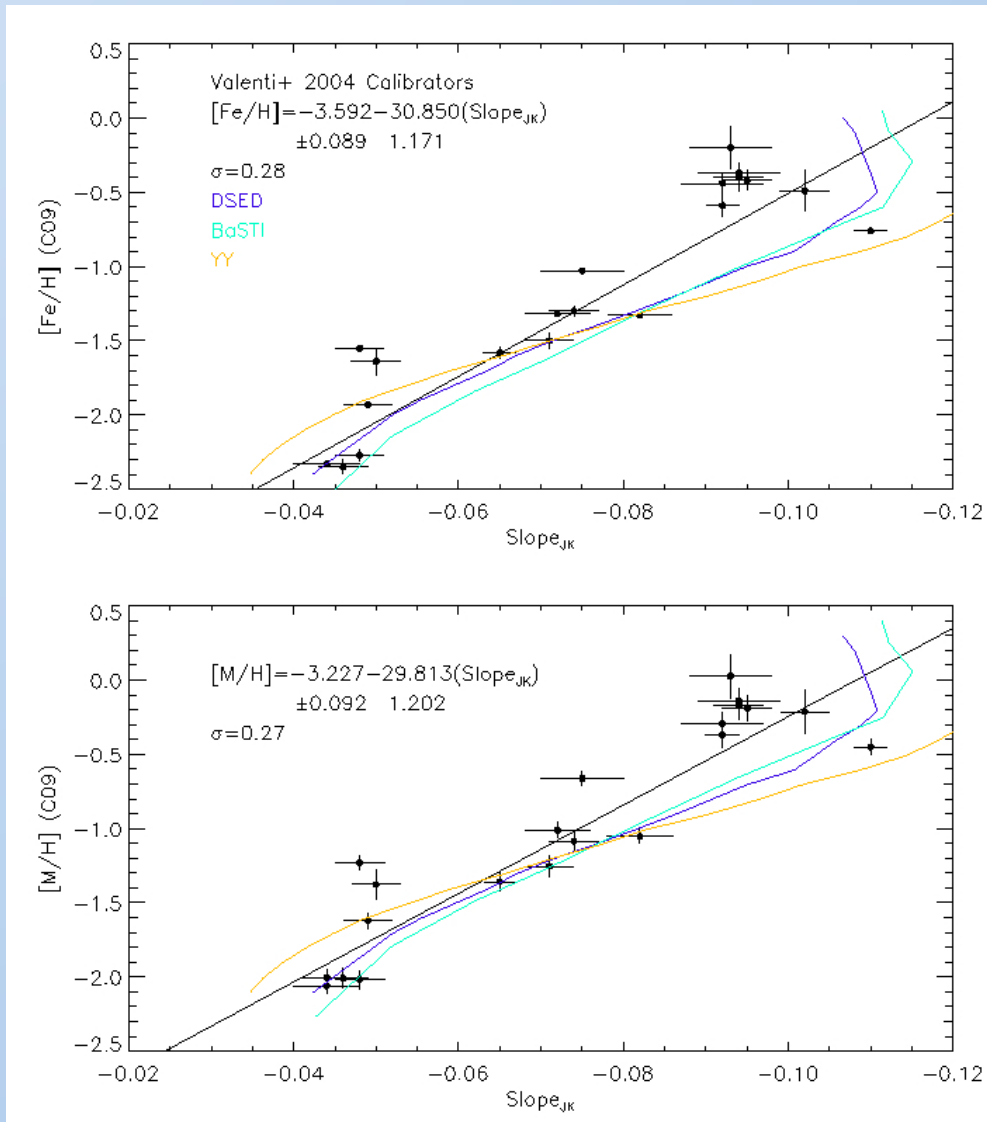
- Known variables removed
Clement+01 update, Samus+09, others
- Reddening maps
Gonzalez+12, Alonso Garcia+12
- Radial cuts
Small scale reddening vs. more cluster stars
- Statistical field star decontamination
Subtraction efficiency > 93% for all clusters
cf. Bonatto et al.

Red Giant Branch Analysis



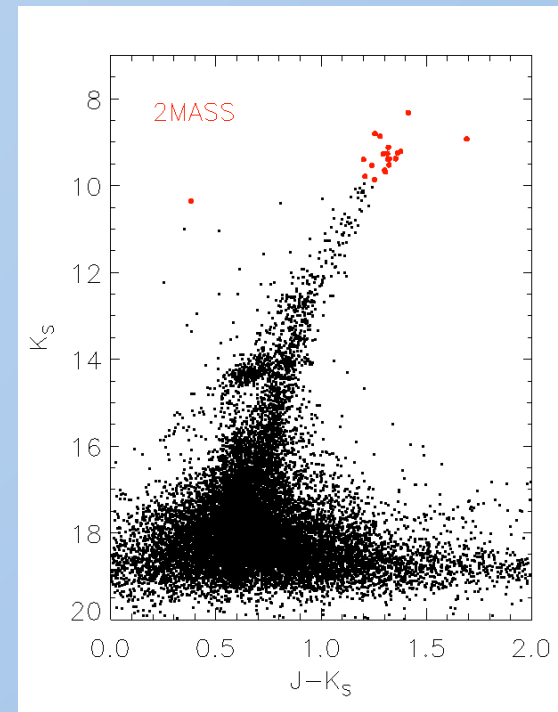
- Known variables removed
Clement+01 update, Samus+09, others
- Reddening maps
Gonzalez+12, Alonso Garcia+12
- Radial cuts
Small scale reddening vs. more cluster stars
- Statistical field star decontamination
Subtraction efficiency > 93% for all clusters
cf. Bonatto et al.

RGB Parameters: Slope

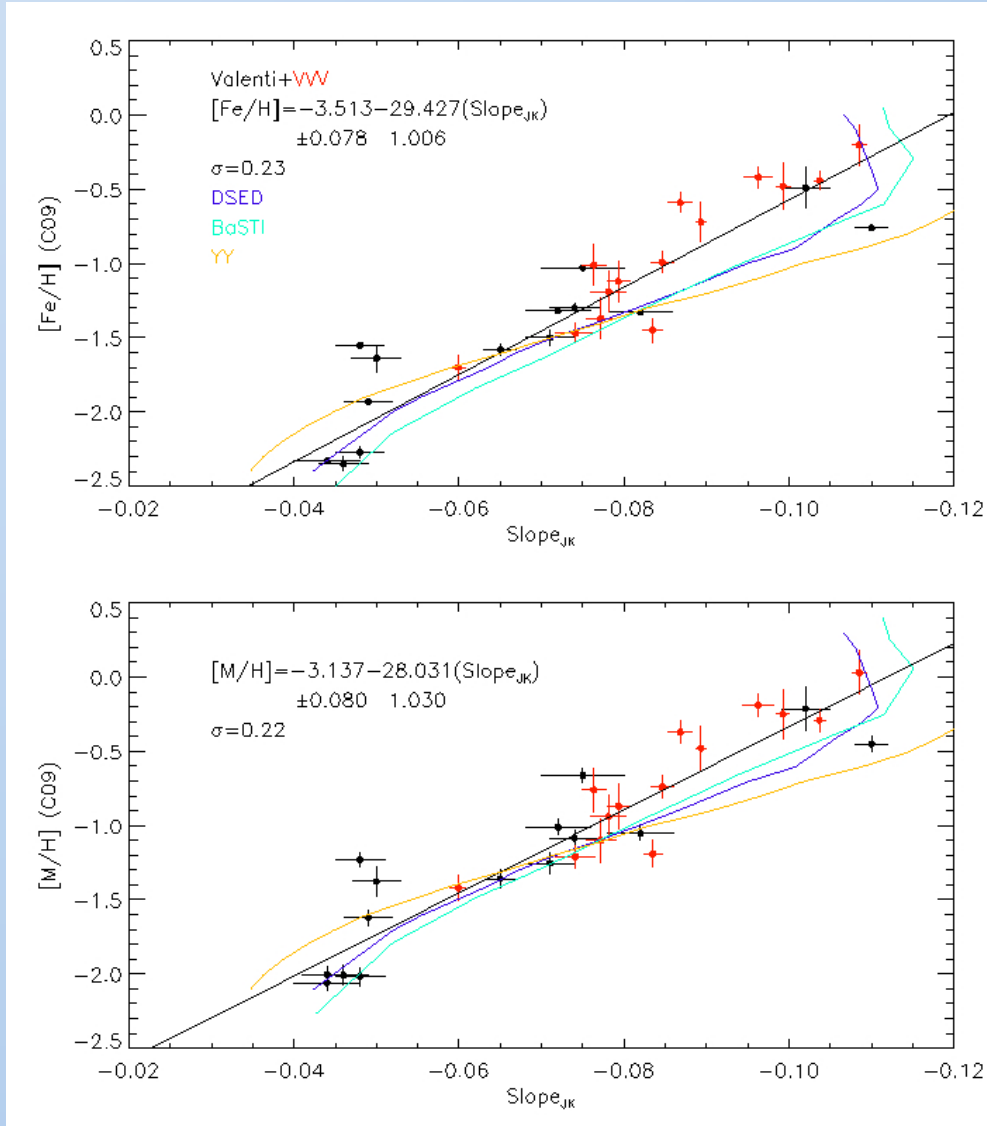


- ~14 GGCs in VVV have $[Fe/H]$ measured spectroscopically

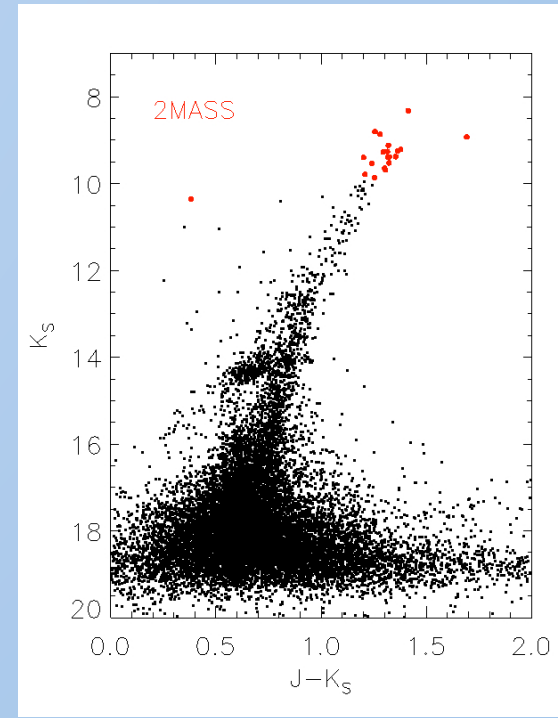
- $K_{tip} + 0.5 < K_s < K_{tip} + 5$ to avoid HB, RGB, AGB variables



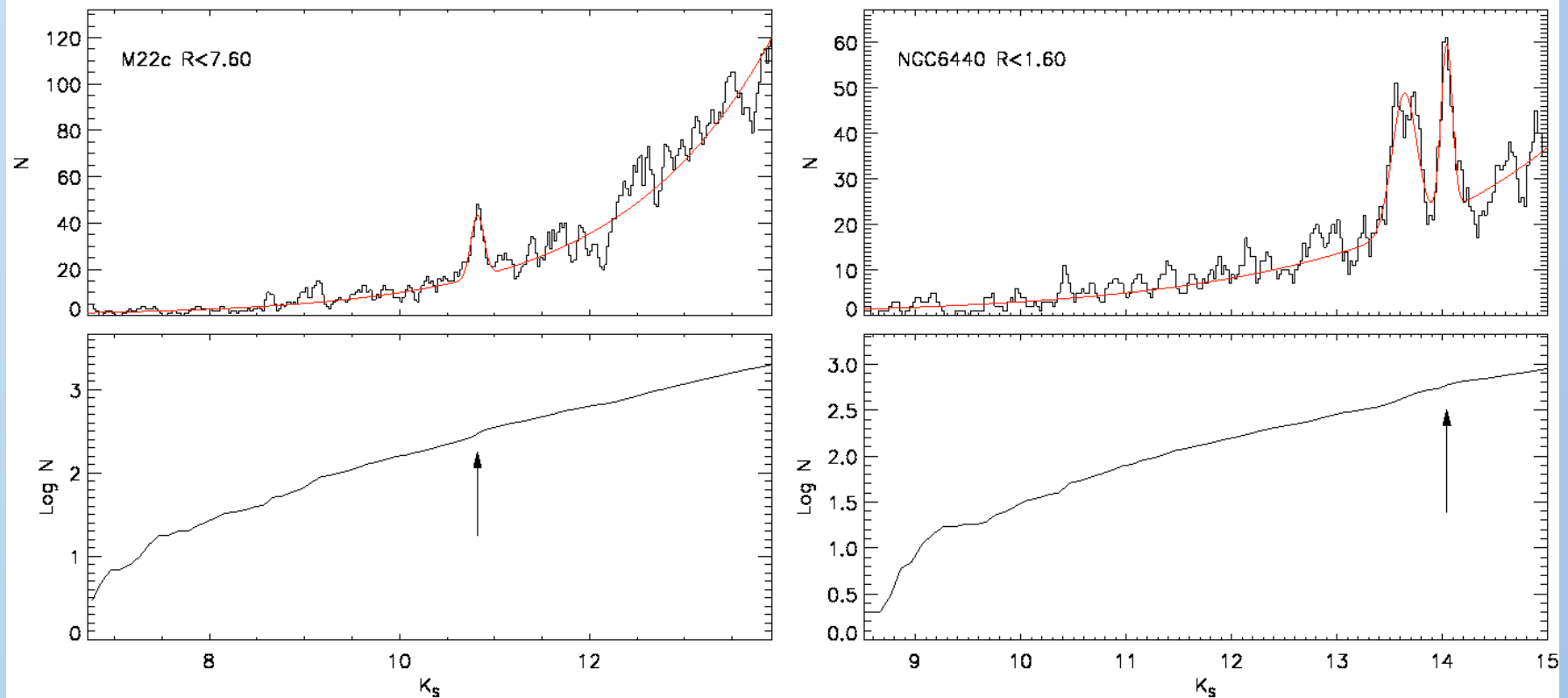
RGB Parameters: Slope



- ~14 GGCs in VVV have $[Fe/H]$ measured spectroscopically
- $K_{tip} + 0.5 < K_s < K_{tip} + 5$ to avoid HB, RGB, AGB variables

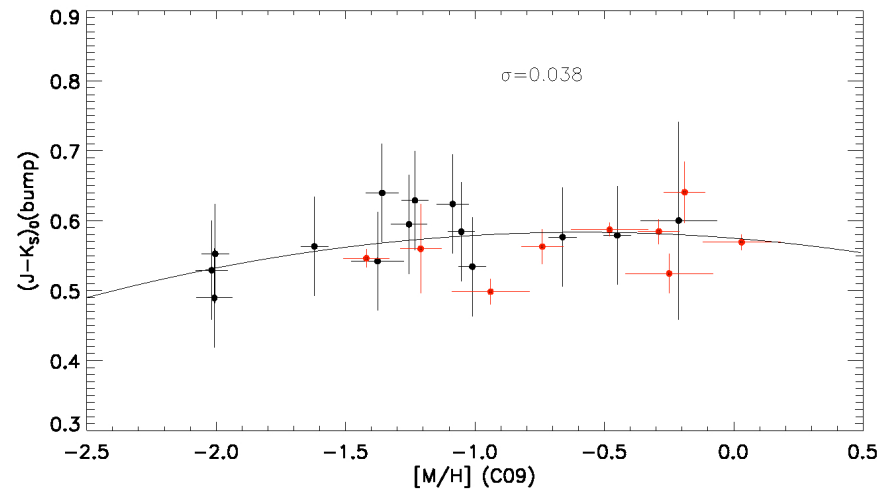
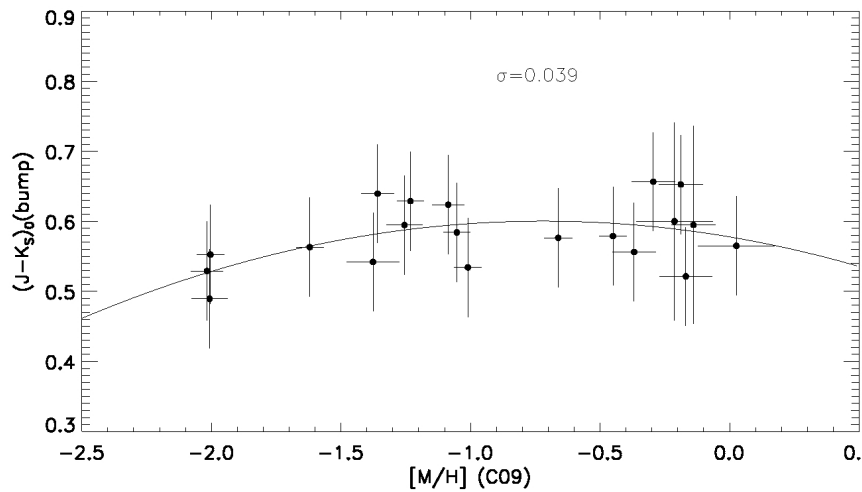
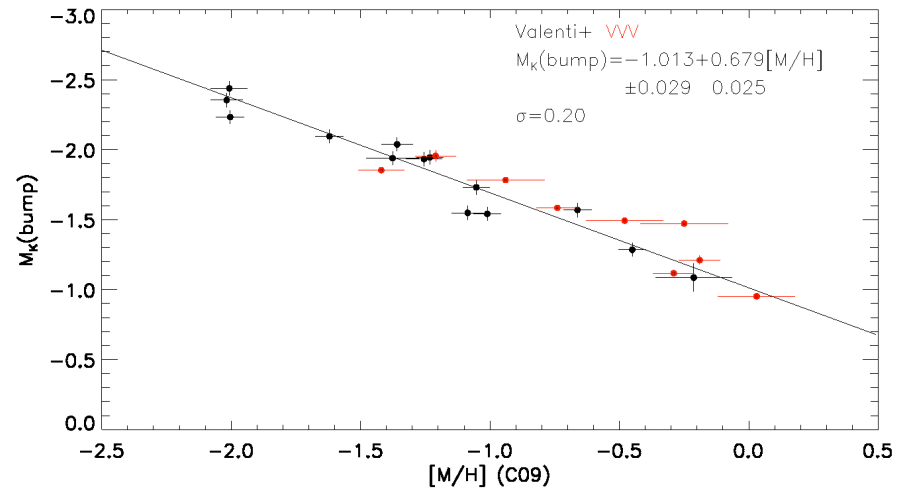
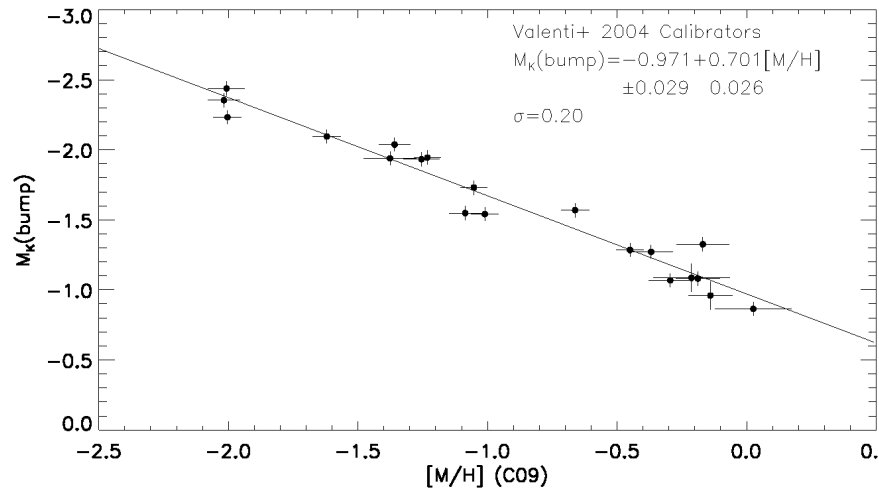


RGB Parameters: Luminosity Function

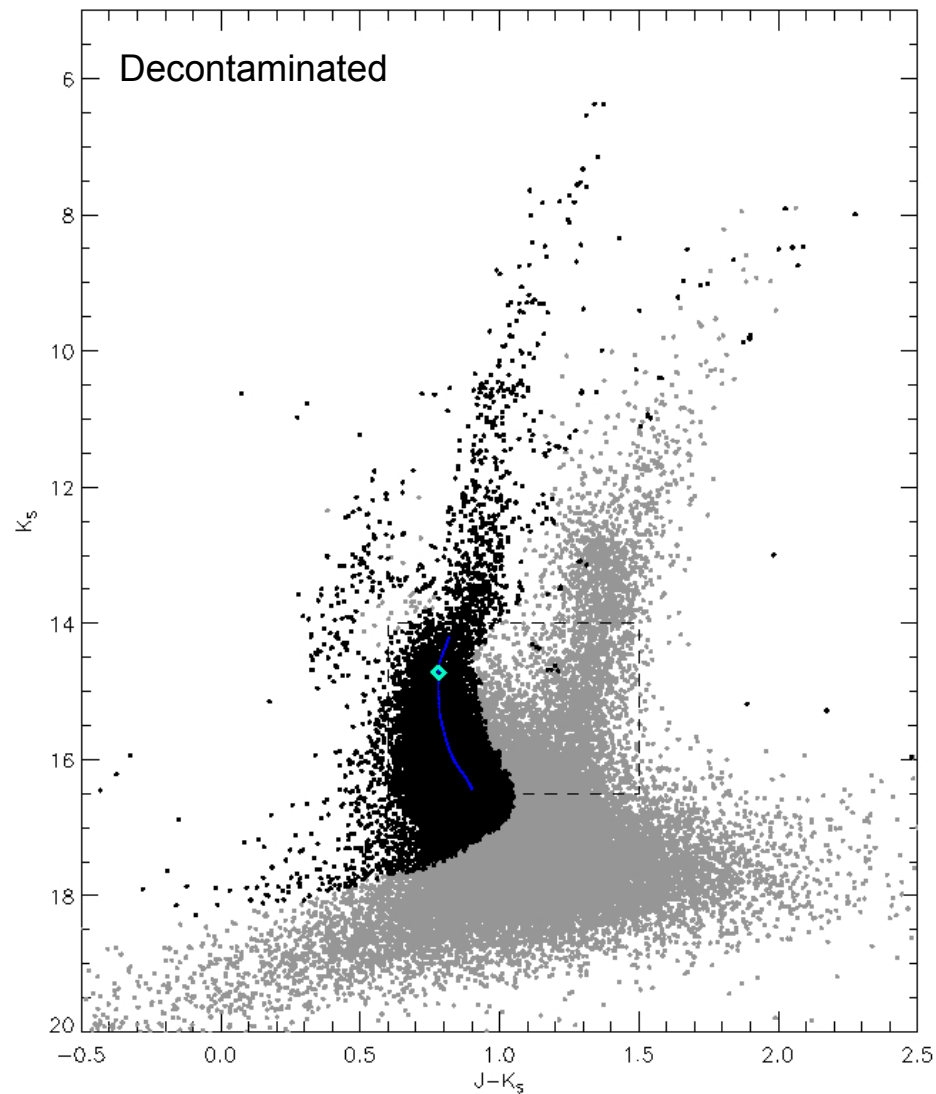


Fit with exponential + (double) Gaussian (cf. Nataf et al. 2013)

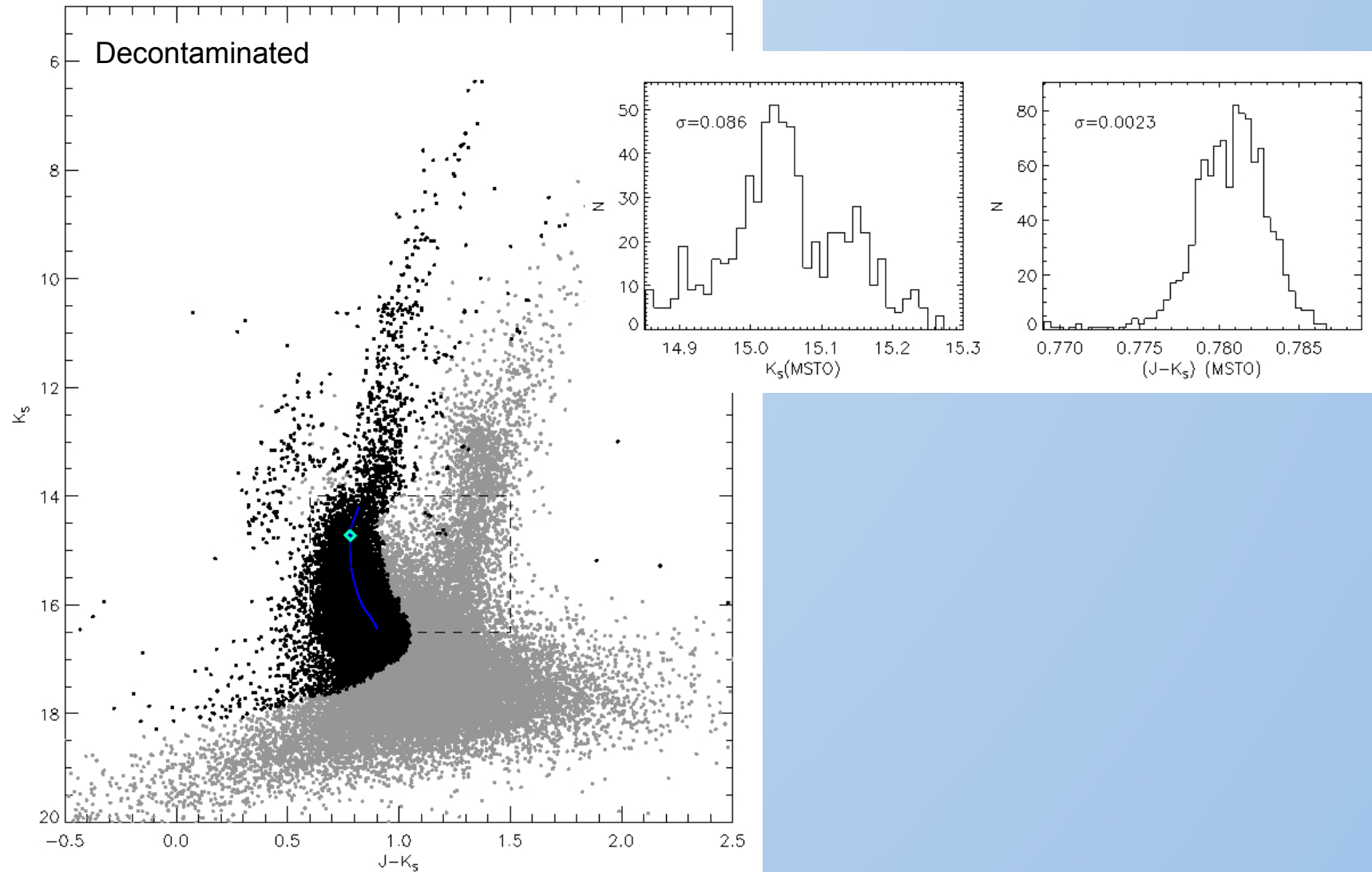
RGB Parameters: RGB Bump



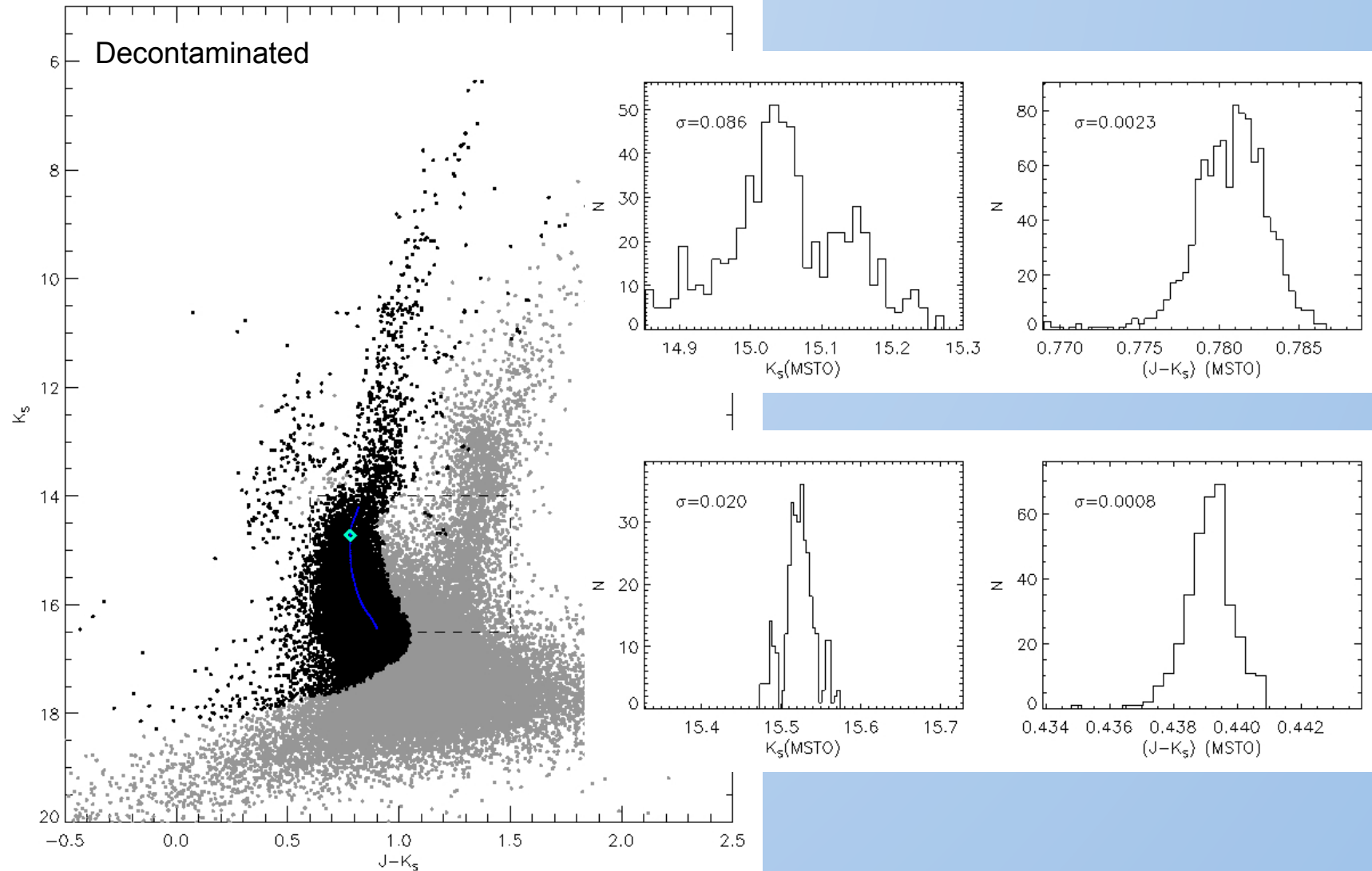
Ages of Metal-Poor Globular Clusters in the Near IR



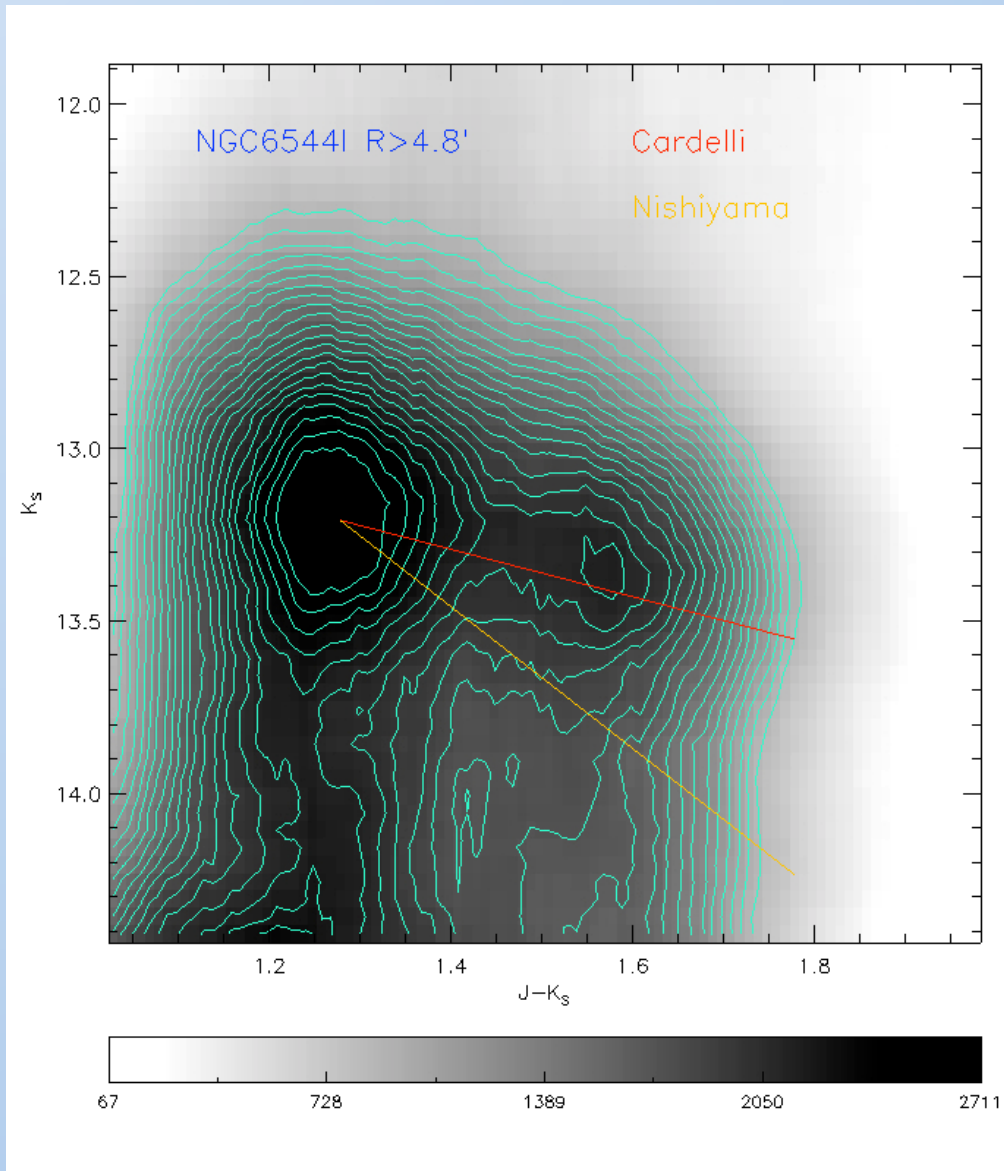
Ages of Metal-Poor Globular Clusters in the Near IR



Ages of Metal-Poor GGCs in the Near IR



Distances and the Reddening Law



Cardelli et al. (1989):

Optical: $R_V = 3.1$

IR: $A(\lambda) \sim \lambda^{-\alpha}$, $\alpha = 1.61$

Not true in the bulge!

Optical: $R_V \sim 2.5$ (Nataf+12)

“non standard and non uniform”

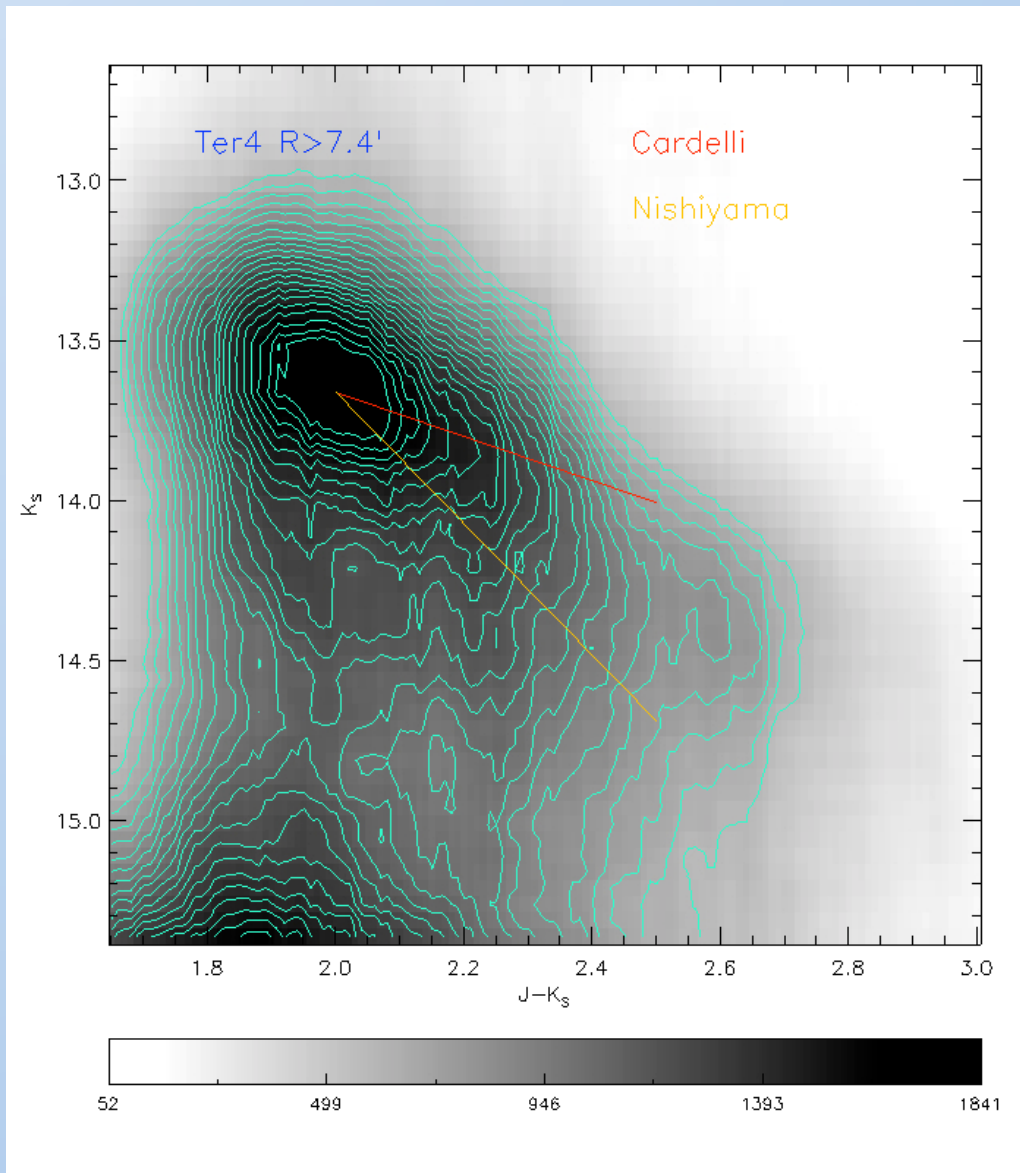
IR: $\alpha \sim 2.0$ (Nishiyama+06)

for $|l| < 2$, $0.5 < |b| < 1$

$\Delta R_V = 0.6$ for $E(B-V) = 1$:

$\Delta d \sim 2$ kpc near bulge!

Distances and the Reddening Law



Cardelli et al. (1989):

Optical: $R_V=3.1$

IR: $A(\lambda) \sim \lambda^{-\alpha}$, $\alpha=1.61$

Not true in the bulge!

Optical: $R_V \sim 2.5$ (Nataf+12)

“non standard and non uniform”

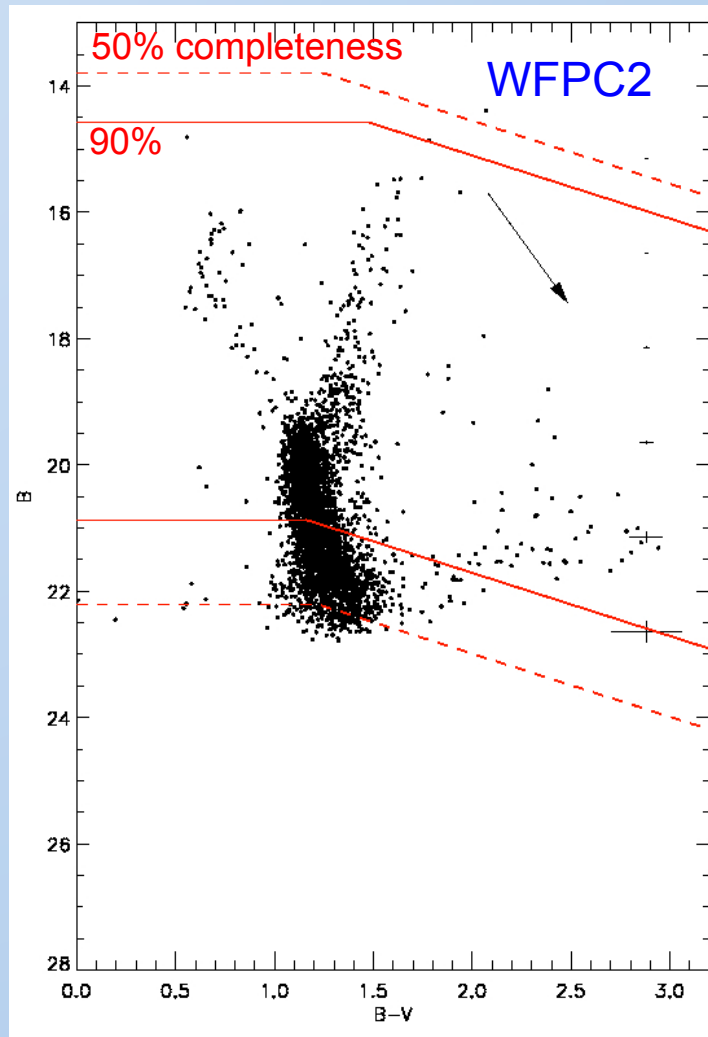
IR: $\alpha \sim 2.0$ (Nishiyama+06)

for $|l| < 2$, $0.5 < |b| < 1$

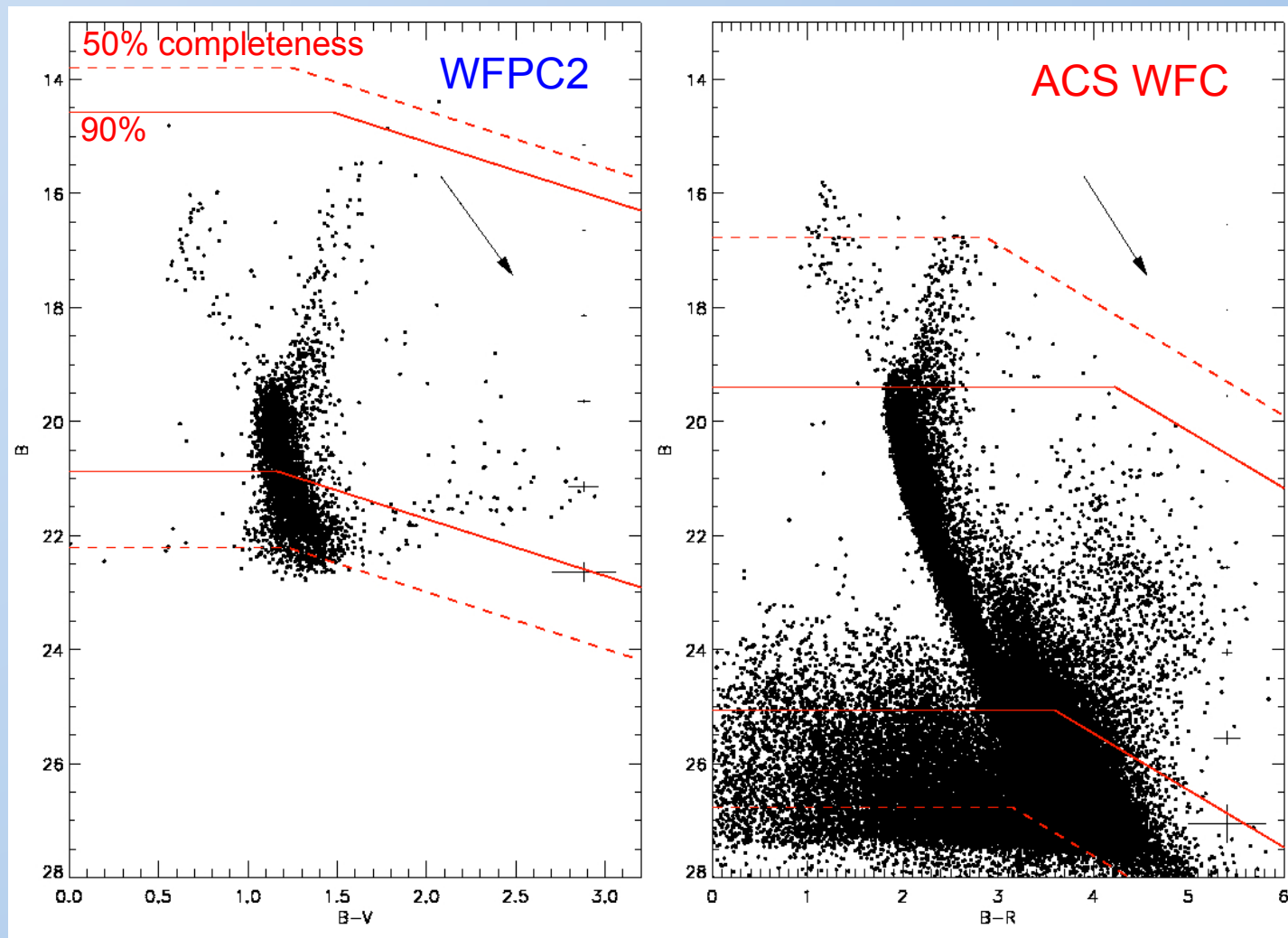
$\Delta R_V = 0.6$ for $E(B-V)=1$:

$\Delta d \sim 2$ kpc near bulge!

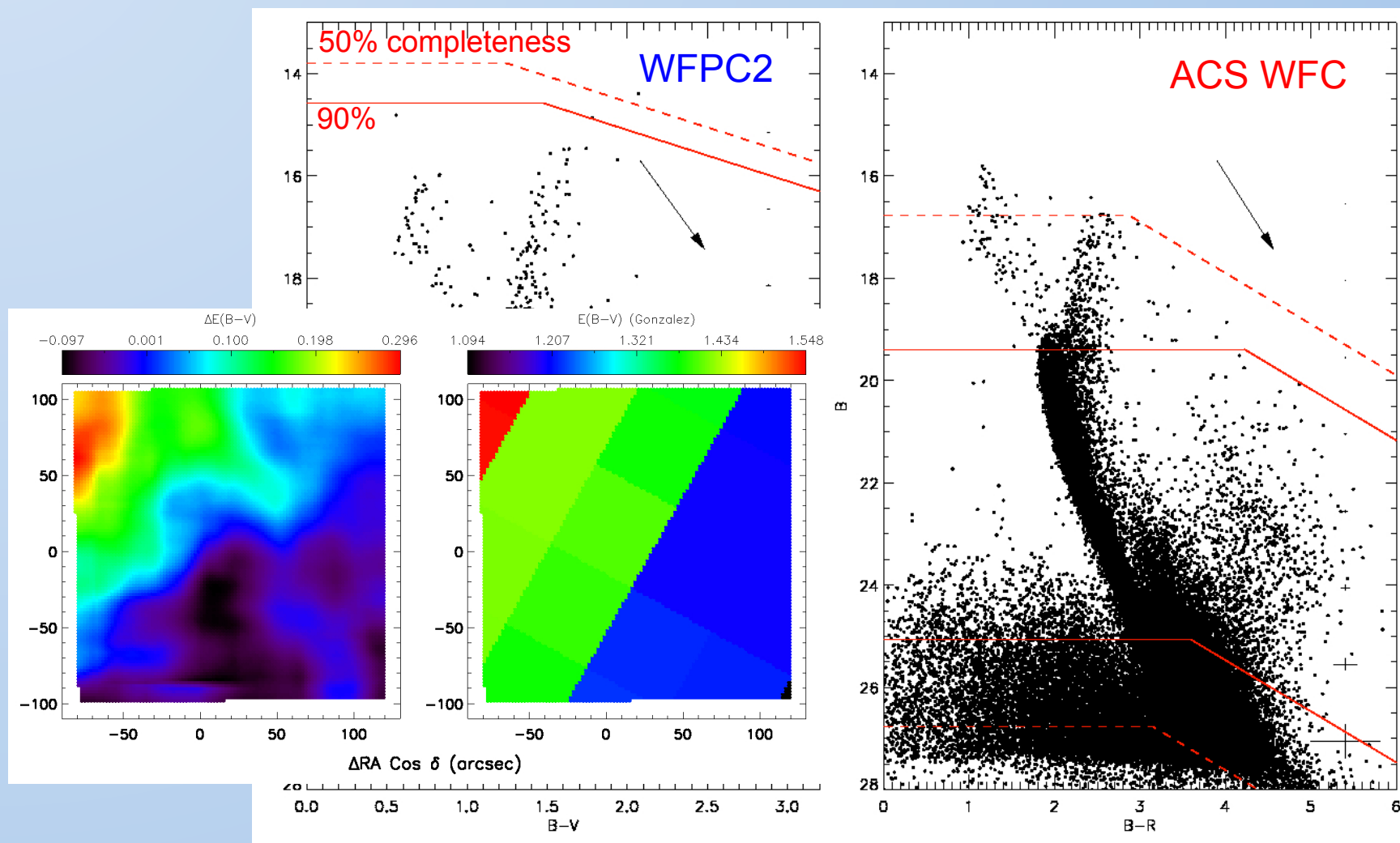
VVV + Optical Photometry



VVV + Optical Photometry

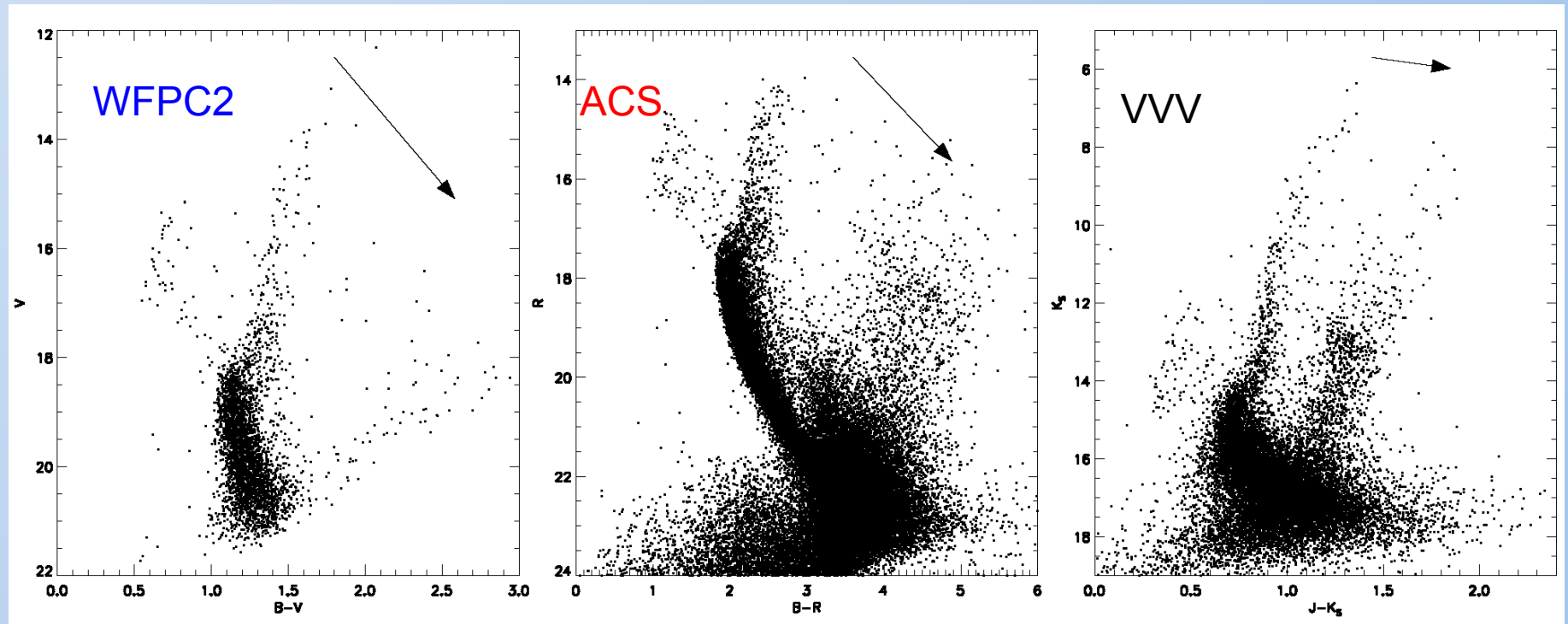


VVV + Optical Photometry

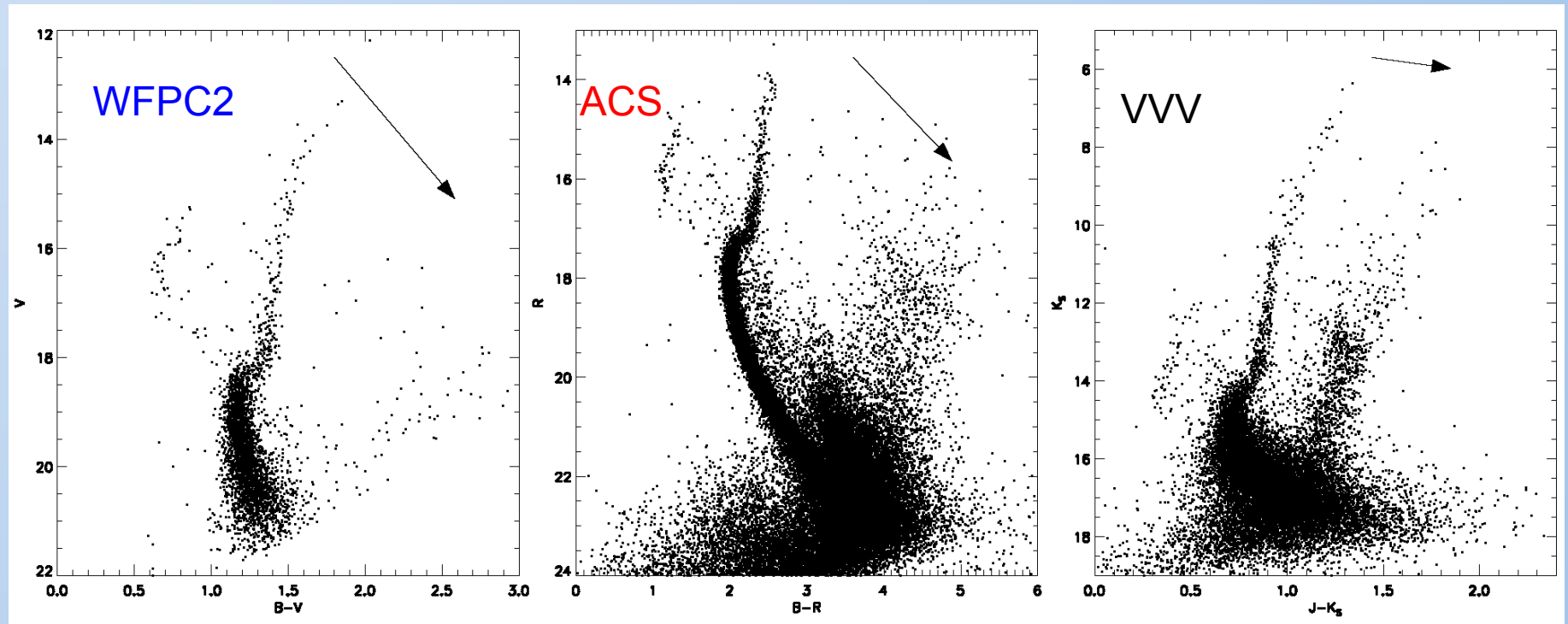


Exploit reddening sensitivity of deep optical photometry (e.g. Alonso Garcia+11,12)

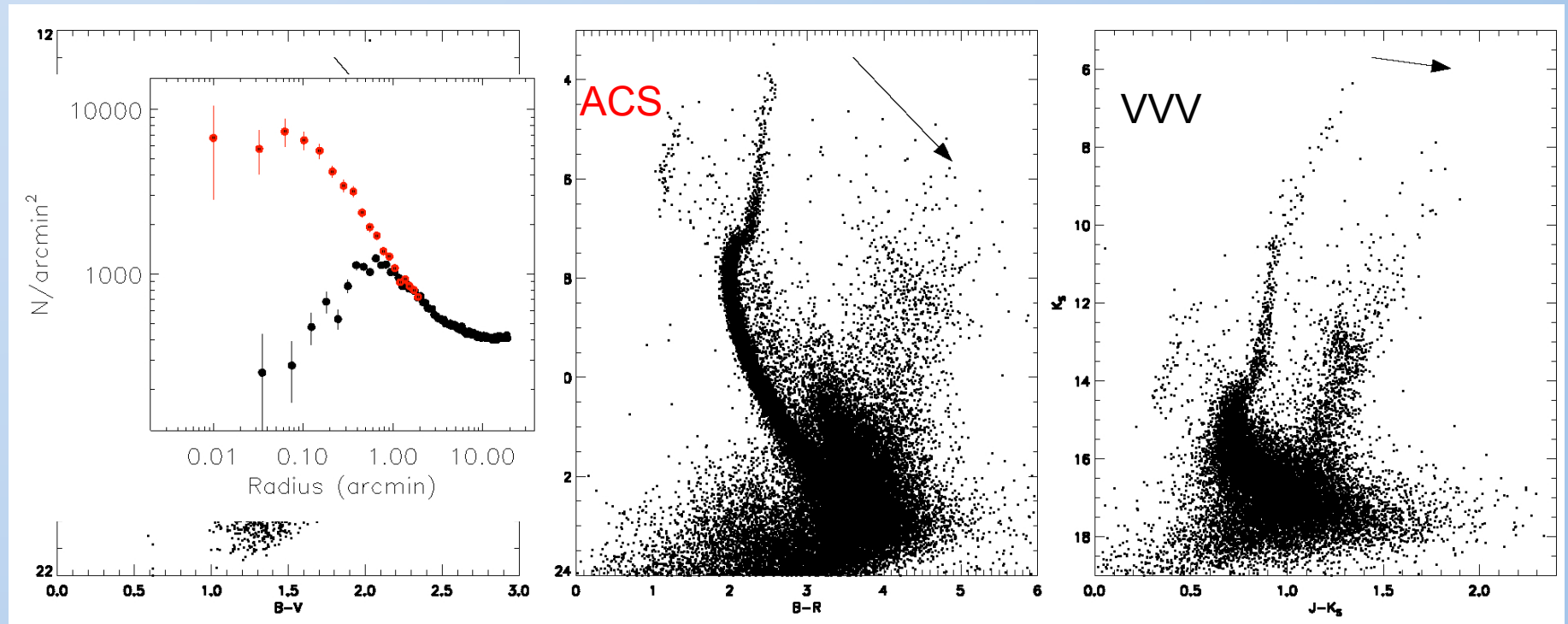
VVV + Optical Photometry



VVV + Optical Photometry



VVV + Optical Photometry



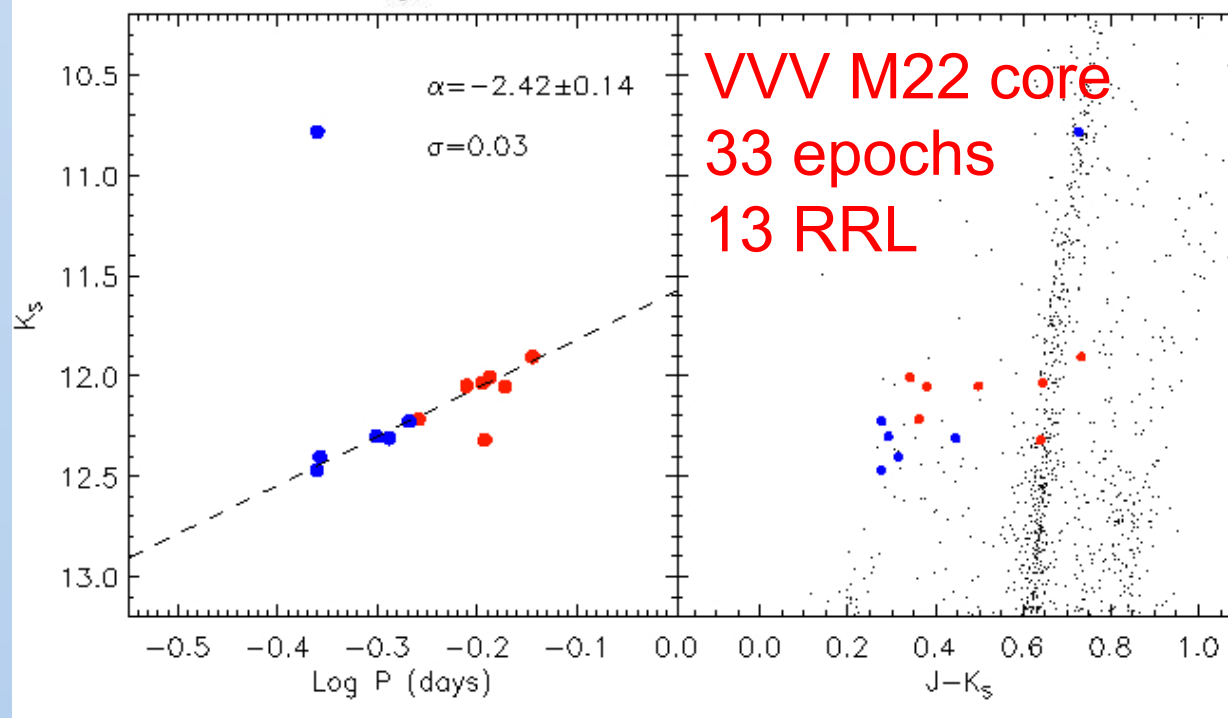
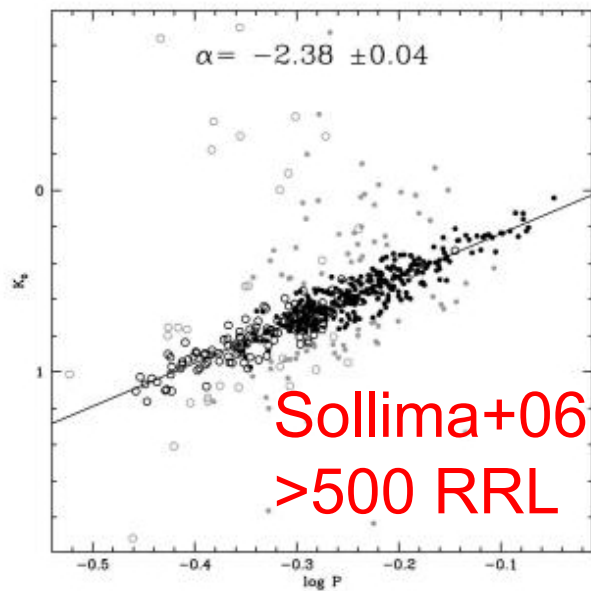
Summary and Future Plans

- JHKs PSF photometry catalogs for all clusters VVV-SkZ pipeline, Mauro+13
- Follow-up spectroscopy, photometry (deep AO)
- RGB calibrations: Apply to clusters without spectroscopic abundances
- Structural parameters
- Optical-IR calibrations

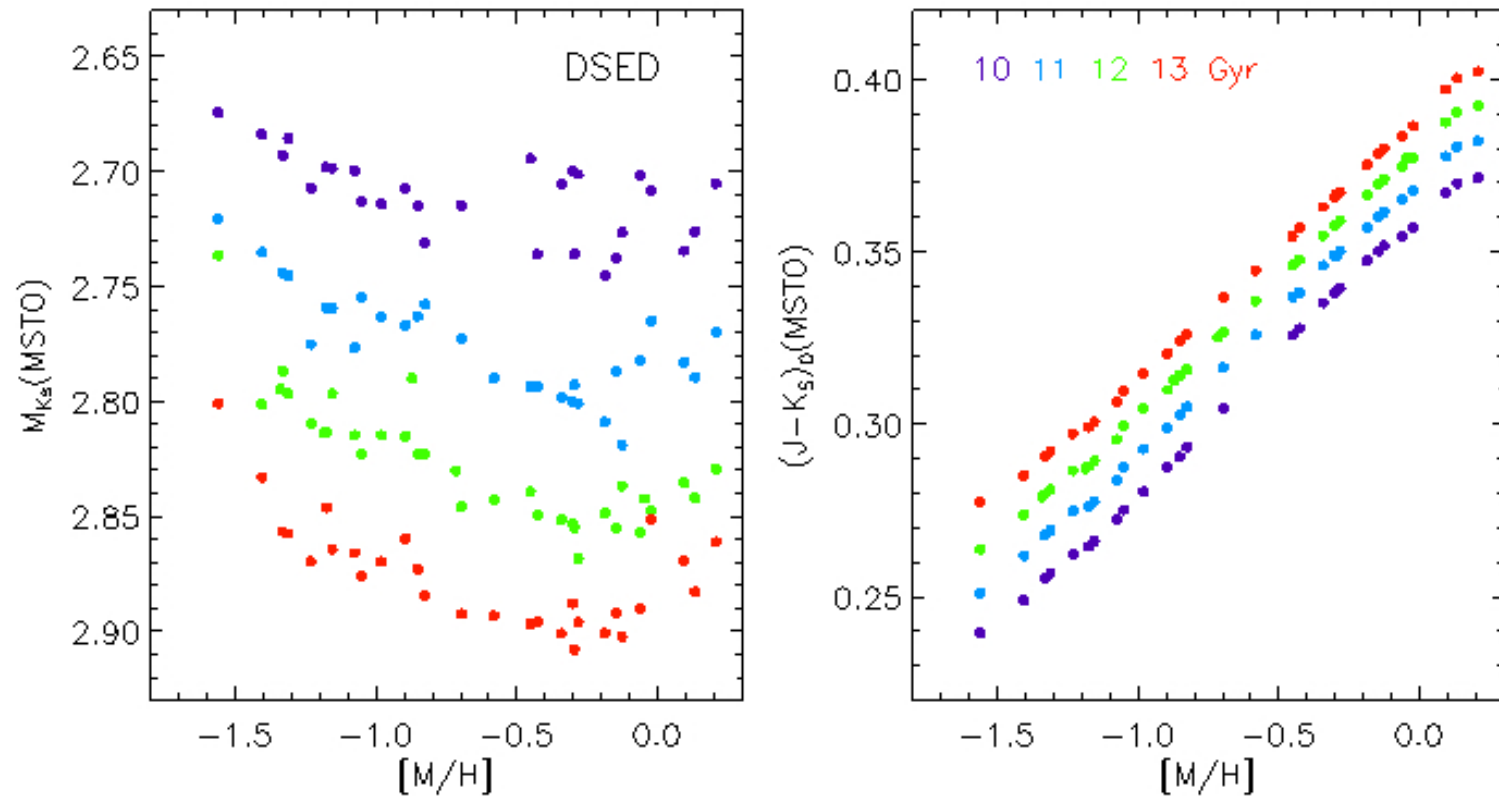
Long term:

- Variability, period-luminosity relations (J. Alonso et al.)
- Relative proper motions (J. C. Beamin et al.)

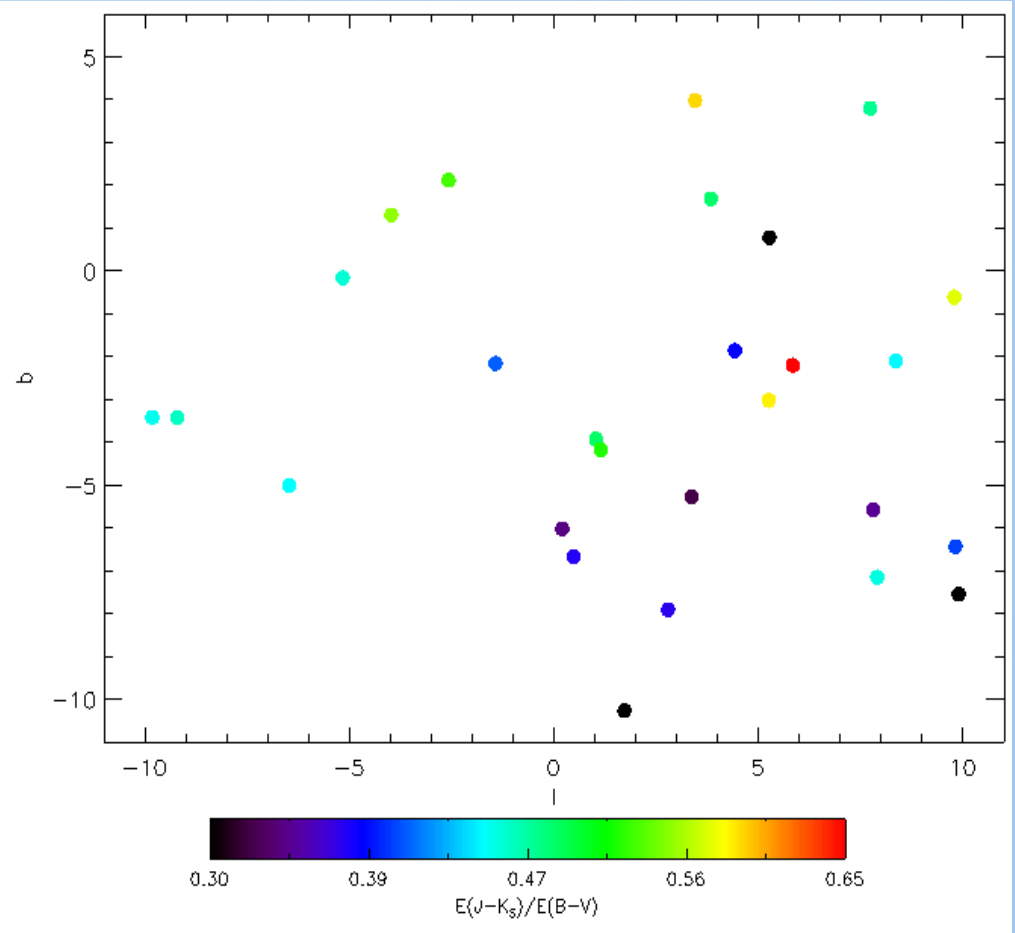
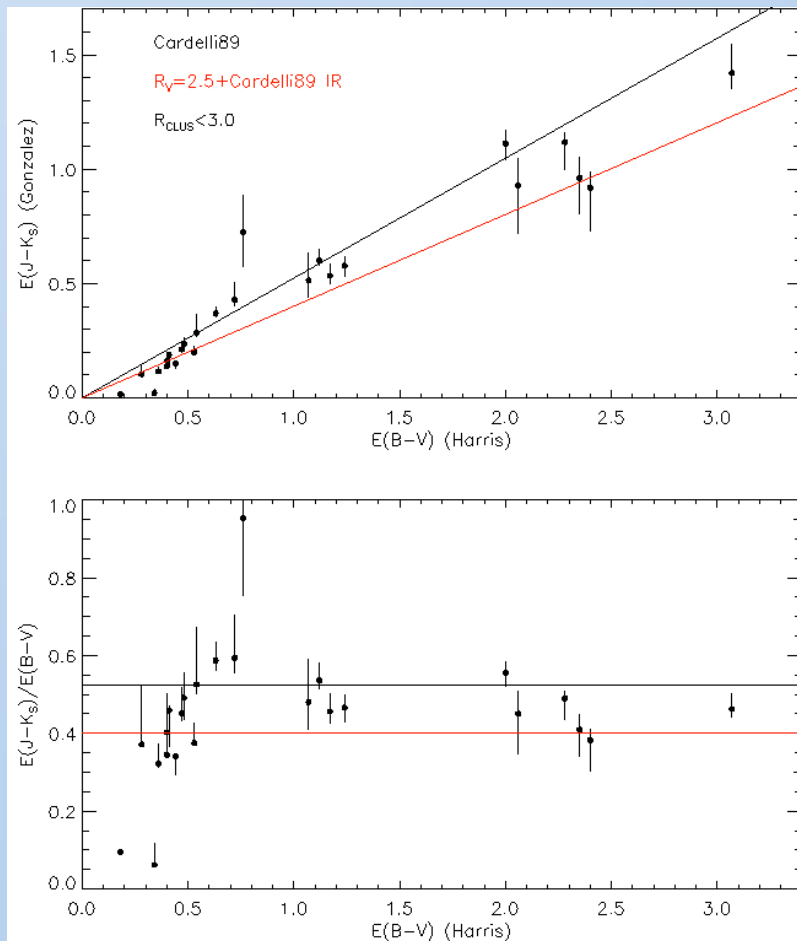
Globular cluster RRL period-luminosity relation with VVV Ks time series



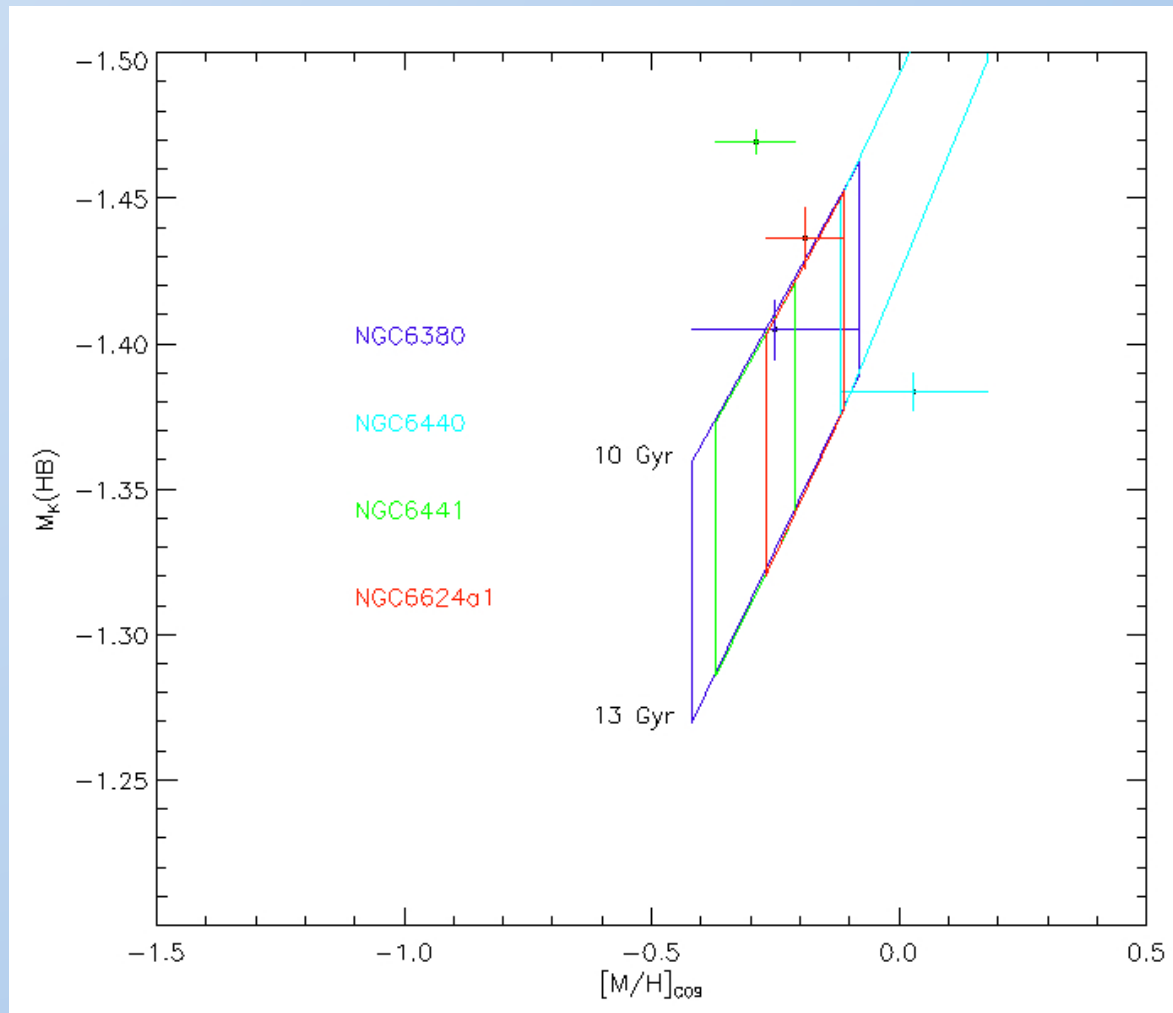
MSTO Color, Mag vs. Age



Optical vs. IR Reddening



Red Clump Magnitude



M22 Turnoff

